

# **Neutrino Experiments with Nuclear Emulsion**

**- OPERA & J-PARC T60 -**

福田 努 (東邦大学)

Tsutomu Fukuda (Toho Univ.)



# Result from OPERA

<b>Belgium</b> ULB Brussels		<b>Italy</b> Bari Bologna LNF Frascati L'Aquila LNGS Naples Padova Rome Salerno		<b>Korea</b> Jinju	
<b>Croatia</b> IRB Zagreb				<b>Russia</b> INR RAS Moscow LPI RAS Moscow ITEP Moscow SINP MSU Moscow JINR Dubna	
<b>France</b> LAPP Annecy IPHC Strasbourg				<b>Switzerland</b> Bern	
<b>Germany</b> Hamburg		<b>Japan</b> Aichi edu. Kobe Nagoya Toho Nihon		<b>Turkey</b> METU Ankara	
<b>Israel</b> Technion Haifa					





## Discovery of $\tau$ Neutrino Appearance in the CNGS Neutrino Beam with the OPERA Experiment

N. Agafonova,<sup>1</sup> A. Aleksandrov,<sup>2</sup> A. Anokhina,<sup>3</sup> S. Aoki,<sup>4</sup> A. Ariga,<sup>5</sup> T. Ariga,<sup>5</sup> D. Bender,<sup>6</sup> A. Bertolin,<sup>7</sup> I. Bodnarchuk,<sup>8</sup> C. Bozza,<sup>9</sup> R. Brugnera,<sup>7,10</sup> A. Buonaura,<sup>2,11</sup> S. Buontempo,<sup>2</sup> B. Büttner,<sup>12</sup> M. Chernyavsky,<sup>13</sup> A. Chukanov,<sup>8</sup> L. Consiglio,<sup>2</sup> N. D'Ambrosio,<sup>14</sup> G. De Lellis,<sup>2,11</sup> M. De Serio,<sup>15,16</sup> P. Del Amo Sanchez,<sup>17</sup> A. Di Crescenzo,<sup>2</sup> D. Di Ferdinando,<sup>18</sup> N. Di Marco,<sup>14</sup> S. Dmitrievski,<sup>8</sup> M. Dracos,<sup>19</sup> D. Duchesneau,<sup>17</sup> S. Dusini,<sup>7</sup> T. Dzhatdoev,<sup>3</sup> J. Ebert,<sup>12</sup> A. Ereditato,<sup>5</sup> R. A. Fini,<sup>16</sup> F. Fornari,<sup>18,20</sup> T. Fukuda,<sup>21</sup> G. Galati,<sup>2,11</sup> A. Garfagnini,<sup>7,10</sup> J. Goldberg,<sup>22</sup> Y. Gornushkin,<sup>8</sup> G. Grella,<sup>9</sup> A. M. Guler,<sup>6</sup> C. Gustavino,<sup>23</sup> C. Hagner,<sup>12</sup> T. Hara,<sup>4</sup> H. Hayakawa,<sup>24</sup> A. Hollnagel,<sup>12</sup> B. Hosseini,<sup>2,11</sup> K. Ishiguro,<sup>24</sup> K. Jakovcic,<sup>25</sup> C. Jollet,<sup>19</sup> C. Kamiscioglu,<sup>6</sup> M. Kamiscioglu,<sup>6</sup> J. H. Kim,<sup>26</sup> S. H. Kim,<sup>26,\*</sup> N. Kitagawa,<sup>24</sup> B. Klicek,<sup>25</sup> K. Kodama,<sup>27</sup> M. Komatsu,<sup>24</sup> U. Kose,<sup>7,†</sup> I. Kreslo,<sup>5</sup> F. Laudisio,<sup>9</sup> A. Lauria,<sup>2,11</sup> A. Ljubicic,<sup>25</sup> A. Longhin,<sup>28</sup> P. F. Loverre,<sup>23,29</sup> A. Malgin,<sup>1</sup> M. Malenica,<sup>25</sup> G. Mandrioli,<sup>18</sup> T. Matsuo,<sup>21</sup> T. Matsushita,<sup>24</sup> V. Matveev,<sup>1</sup> N. Mauri,<sup>18,20</sup> E. Medinaceli,<sup>7,10</sup> A. Meregaglia,<sup>19</sup> S. Mikado,<sup>30</sup> M. Miyanishi,<sup>24</sup> F. Mizutani,<sup>4</sup> P. Monacelli,<sup>23</sup> M. C. Montesi,<sup>2,11</sup> K. Morishima,<sup>24</sup> M. T. Muciaccia,<sup>15,16</sup> N. Naganawa,<sup>24</sup> T. Naka,<sup>24</sup> M. Nakamura,<sup>24</sup> T. Nakano,<sup>24</sup> Y. Nakatsuka,<sup>24</sup> K. Niwa,<sup>24</sup> S. Ogawa,<sup>21</sup> A. Olchevsky,<sup>8</sup> T. Omura,<sup>24</sup> K. Ozaki,<sup>4</sup> A. Paoloni,<sup>28</sup> L. Paparella,<sup>15,16</sup> B. D. Park,<sup>26,‡</sup> I. G. Park,<sup>26</sup> L. Pasqualini,<sup>18,20</sup> A. Pastore,<sup>15</sup> L. Patrizii,<sup>18</sup> H. Pessard,<sup>17</sup> C. Pistillo,<sup>5</sup> D. Podgrudkov,<sup>3</sup> N. Polukhina,<sup>13</sup> M. Pozzato,<sup>18,20</sup> F. Pupilli,<sup>28</sup> M. Roda,<sup>7,10</sup> T. Roganova,<sup>3</sup> H. Rokujo,<sup>24</sup> G. Rosa,<sup>23,29</sup> O. Ryazhskaya,<sup>1</sup> O. Sato,<sup>24,§</sup> A. Sembri,<sup>14</sup> W. Schmidt-Parzefall,<sup>12</sup> I. Shakirianova,<sup>1</sup> T. Shchedrina,<sup>13,11</sup> A. Sheshukov,<sup>8</sup> H. Shibuya,<sup>21</sup> T. Shiraishi,<sup>24</sup> G. Shoziyoev,<sup>3</sup> S. Simone,<sup>15,16</sup> M. Sioli,<sup>18,20</sup> C. Sirignano,<sup>7,10</sup> G. Sirri,<sup>18</sup> A. Sotnikov,<sup>8</sup> M. Spinetti,<sup>28</sup> L. Stanco,<sup>7</sup> N. Starkov,<sup>13</sup> S. M. Stellacci,<sup>9</sup> M. Stipcevic,<sup>25</sup> P. Strolin,<sup>2,11</sup> S. Takahashi,<sup>4</sup> M. Tenti,<sup>18</sup> F. Terranova,<sup>28,31</sup> V. Tioukov,<sup>2</sup> S. Tufanli,<sup>5,||</sup> P. Vilain,<sup>32</sup> M. Vladymyrov,<sup>13,¶</sup> L. Votano,<sup>28</sup> J. L. Vuilleumier,<sup>5</sup> G. Wilquet,<sup>32</sup> B. Wonsak,<sup>12</sup> C. S. Yoon,<sup>26</sup> and S. Zemskova<sup>8</sup>

(OPERA Collaboration)

arXiv:1507.01417



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The OPERA experiment was designed to search for  $\nu_\mu \rightarrow \nu_\tau$  oscillations in appearance mode, i.e., by detecting the  $\tau$  leptons produced in charged current  $\nu_\tau$  interactions. The experiment took data from 2008 to 2012 in the CERN Neutrinos to Gran Sasso beam. The observation of the  $\nu_\mu \rightarrow \nu_\tau$  appearance, achieved with four candidate events in a subsample of the data, was previously reported. In this Letter, a fifth  $\nu_\tau$  candidate event, found in an enlarged data sample, is described. Together with a further reduction of the expected background, the candidate events detected so far allow us to assess the discovery of  $\nu_\mu \rightarrow \nu_\tau$  oscillations in appearance mode with a significance larger than **5 $\sigma$** .

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PACS numbers: 14.60.Pq

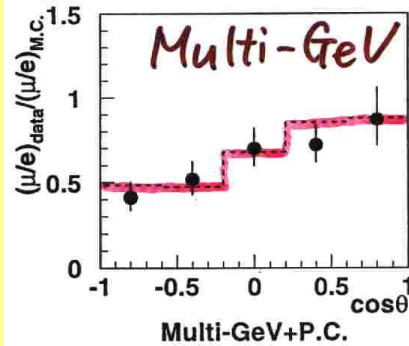
(OPERA Collaboration)

arXiv:1507.01417

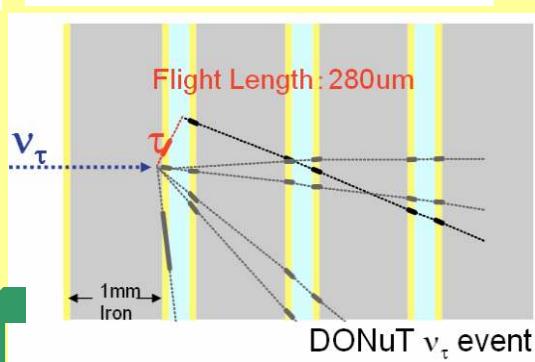
# The OPERA experiment

## Oscillation Project with Emulsion tRacking Apparatus

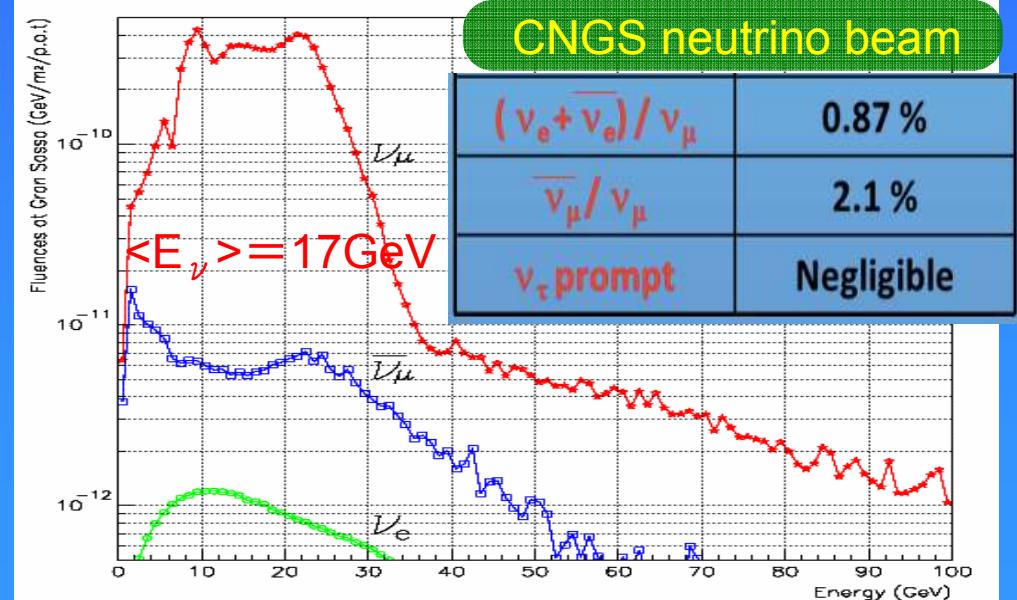
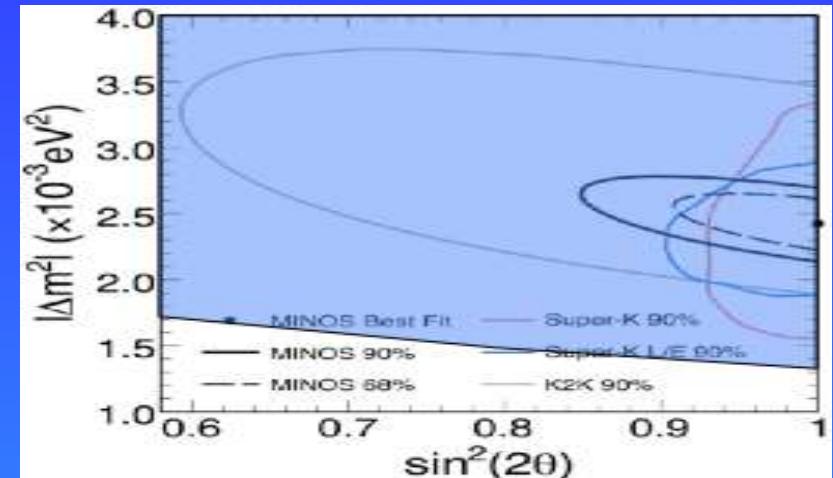
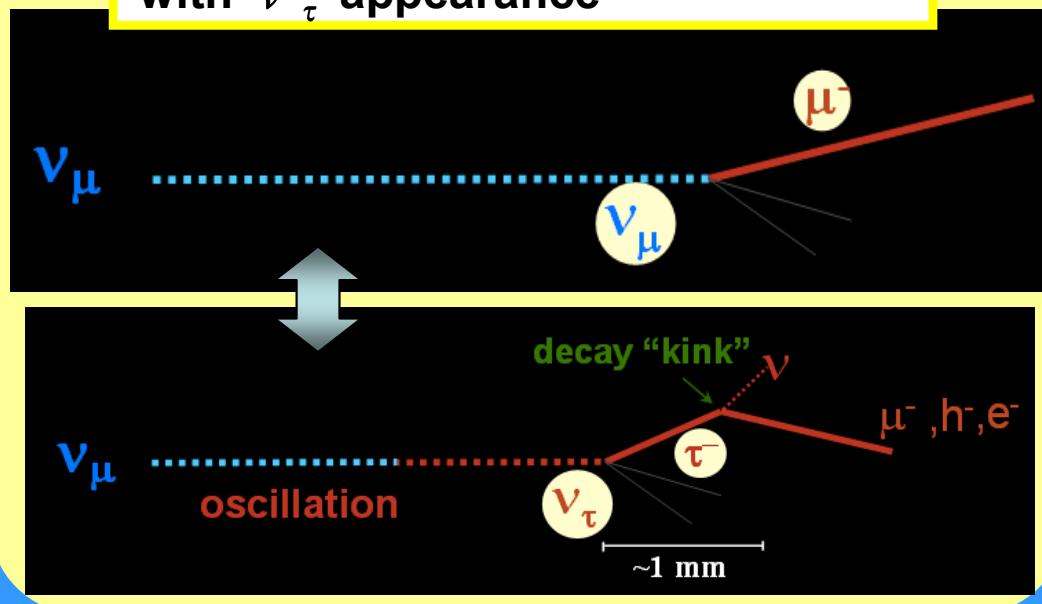
Neutrino oscillation (disappearance)  
Result from SK in 1998



Direct observation of  $\nu_\tau$  events  
Result of DONuT in 2001



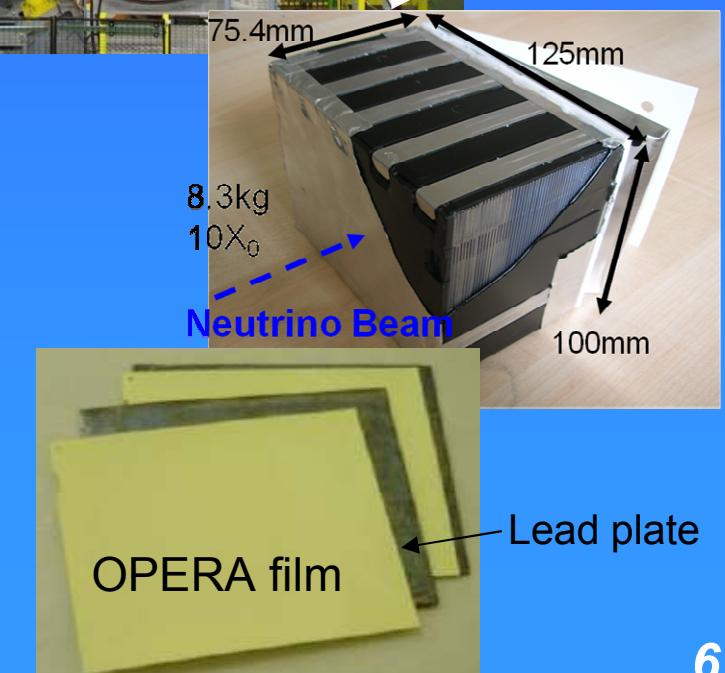
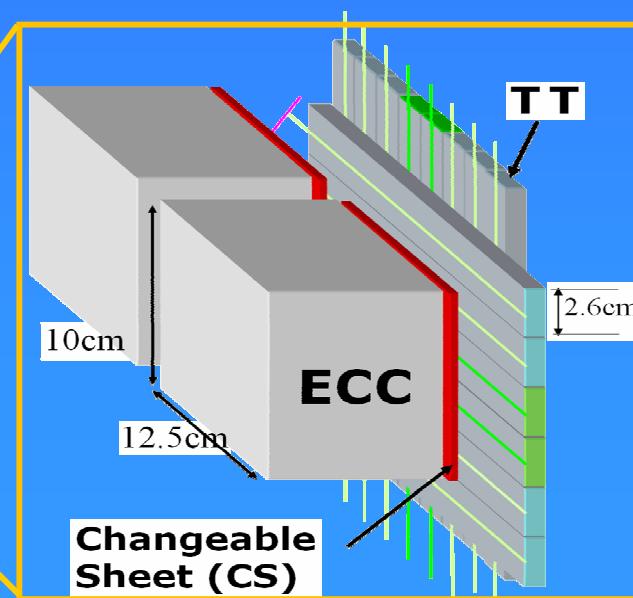
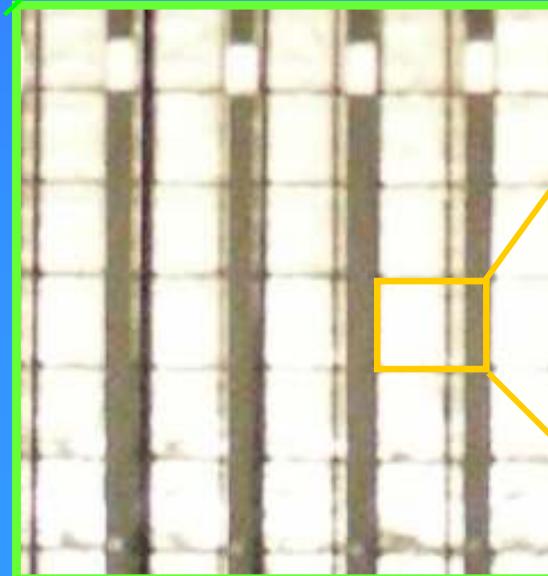
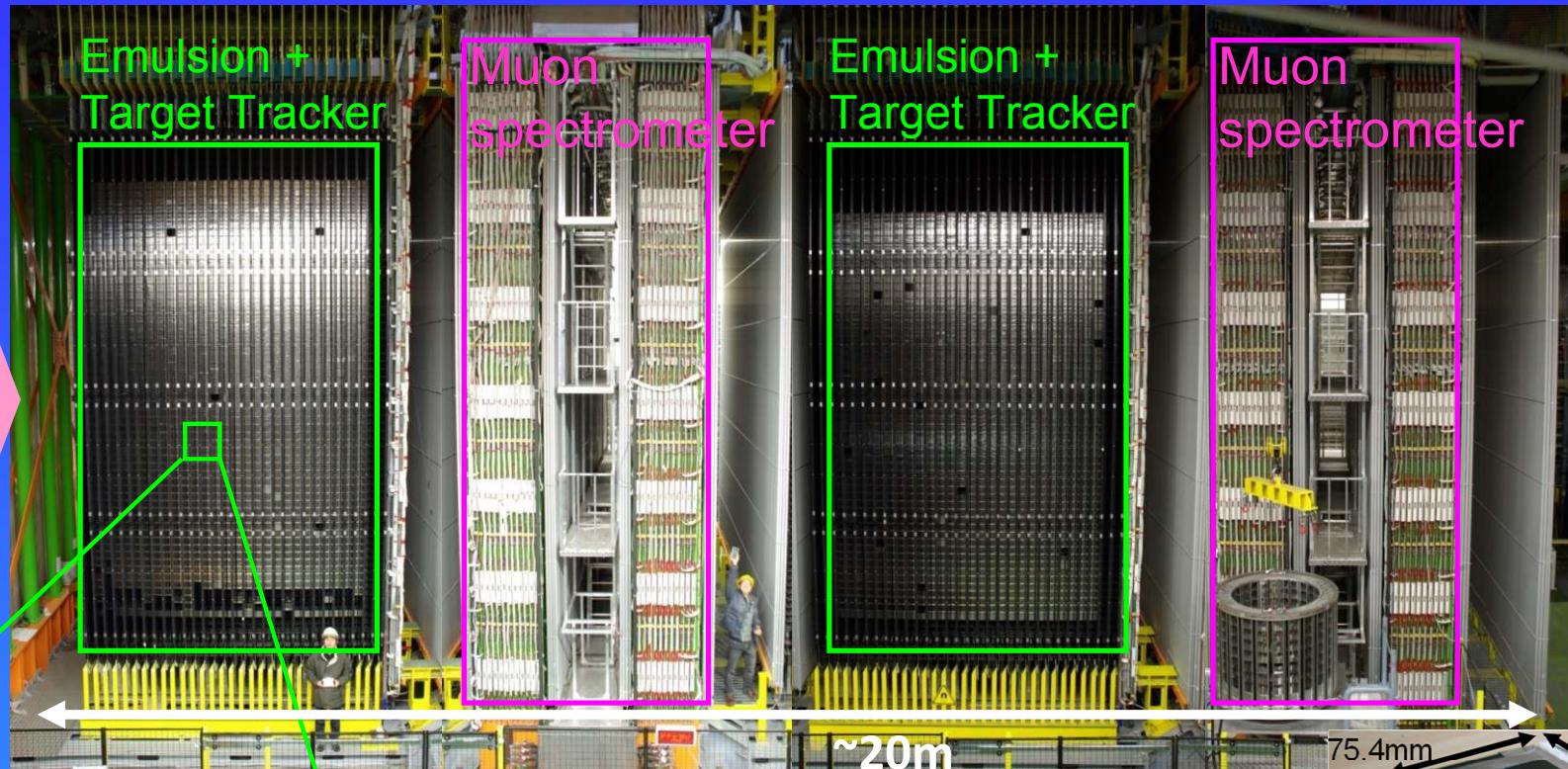
Verification of neutrino oscillation  
with  $\nu_\tau$  appearance



$$P(\nu_\mu \rightarrow \nu_\tau) \sim \sin^2(2\theta_{23}) \cdot \sin^2\left(1.27 \cdot \Delta m_{23}^2 \cdot \frac{L}{E}\right) \sim 1.7\%$$

$$\sin^2 2\theta_{23} = 1.0, \quad \Delta m_{23}^2 = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$$

# The OPERA Detector



# Refresh 900万枚のフィルムのノイズ消し

Facility 建設@東濃鉱山  
2003.8～



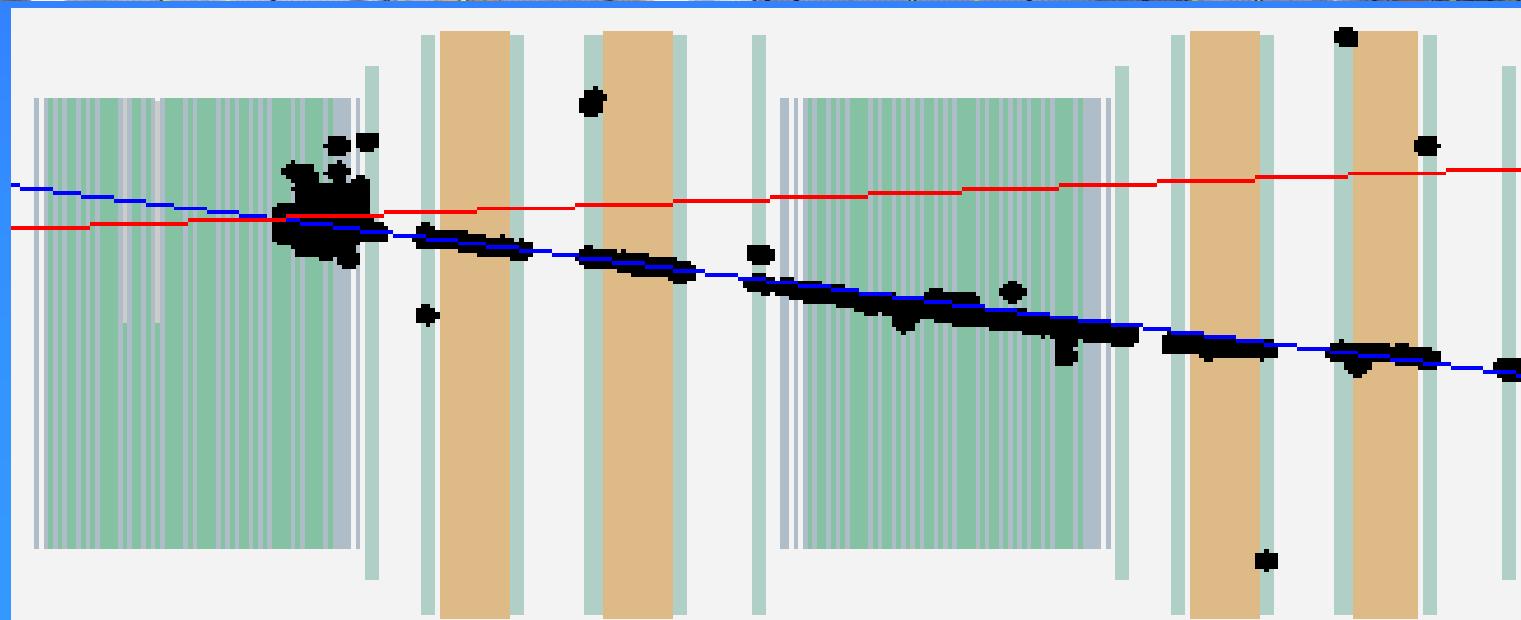
大量処理  
2004.4～



作業ルーチン化R&D  
2003.12～

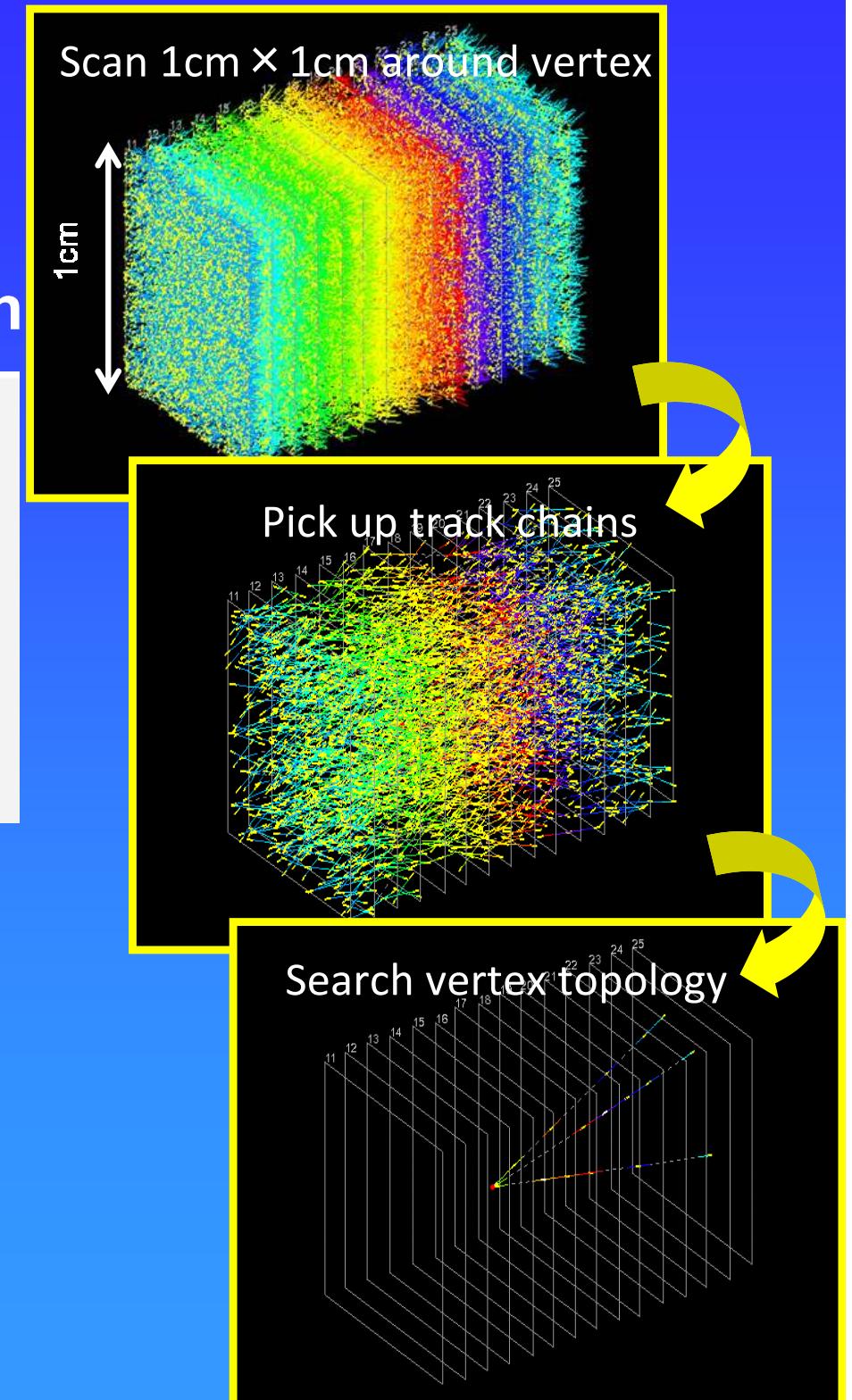
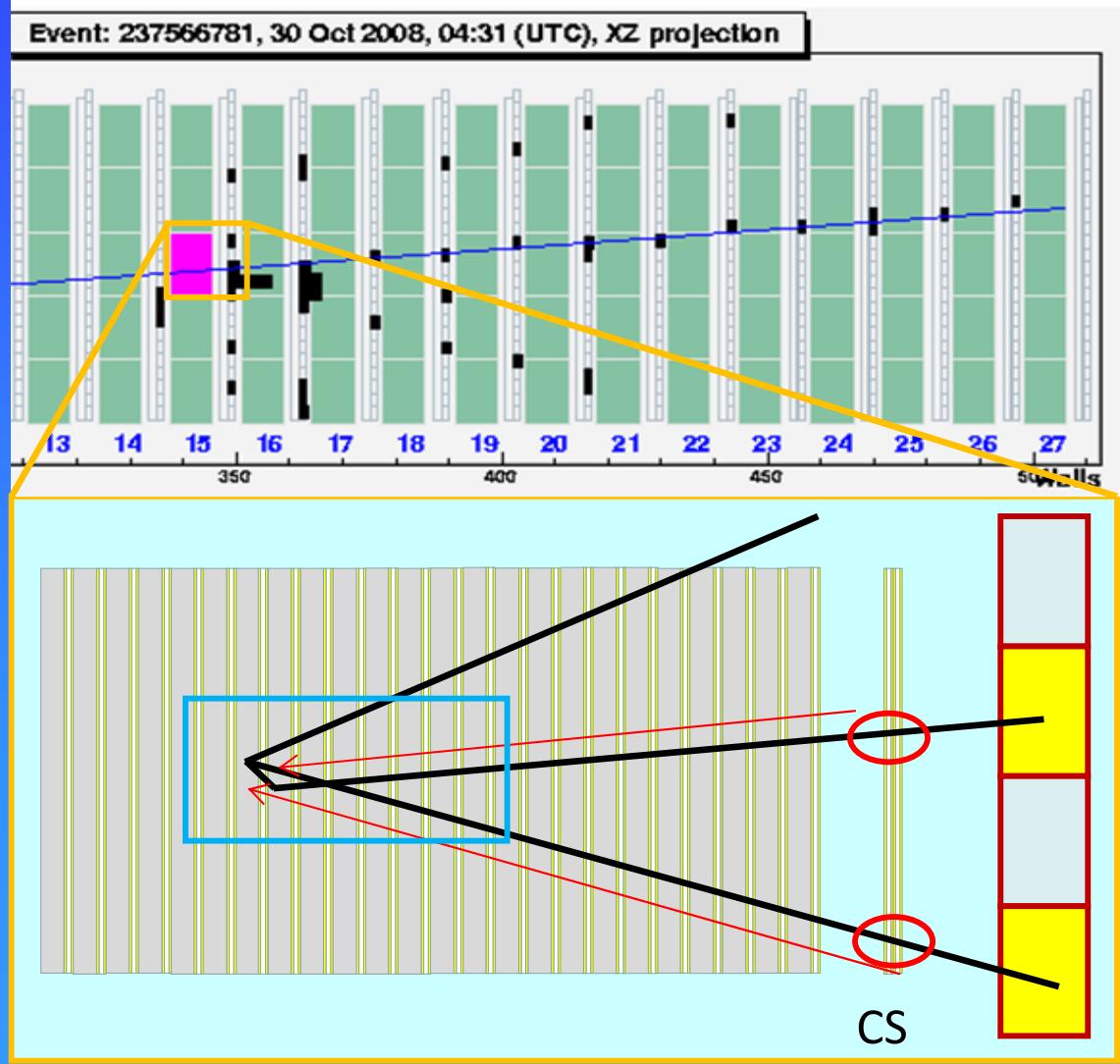


# The OPERA Detector



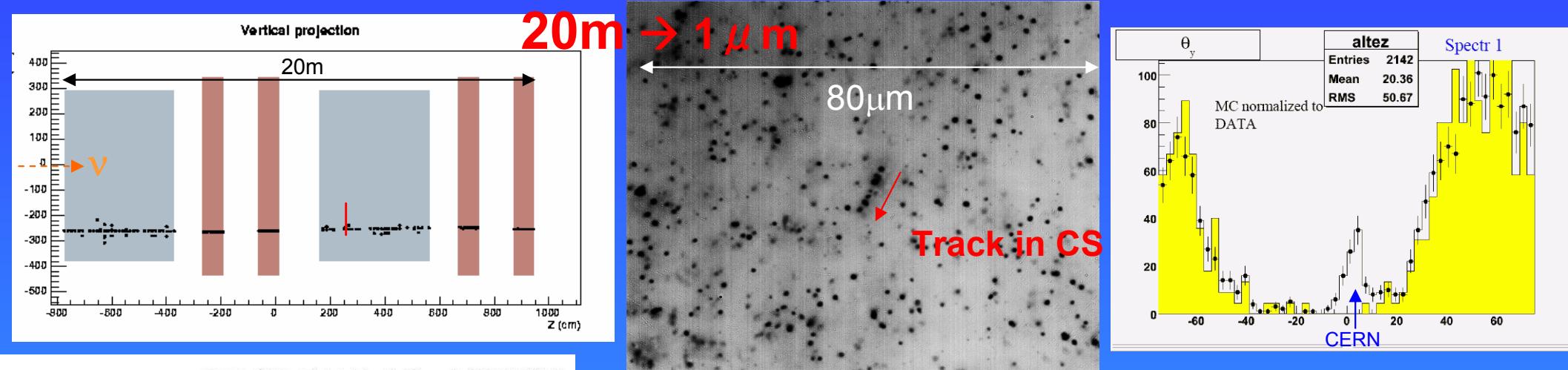
# *Neutrino event analysis*

## Scan Back location & Decay search



# Commissioning-2006

CERNからニュートリノビームをテスト照射。  
OPERAで初めて原子核乾板上にニュートリノ反応からの飛跡を検出。



2006年9月12日(火) 中日新聞朝刊



New J. Phys. 8 (2006) 303

'06 9/7

10

# *Technical improvements*

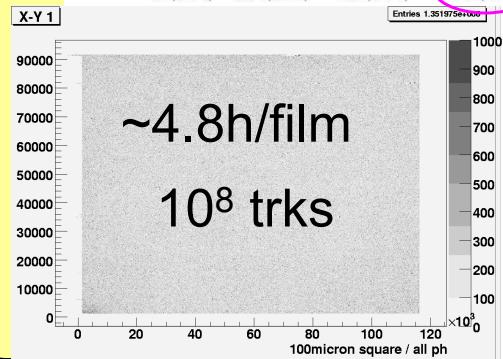
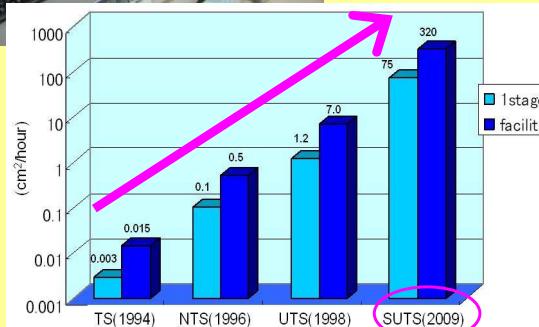
- Challenge in OPERA → 大量のフィルム解析

## SUTS

High speed scanning  
x 60

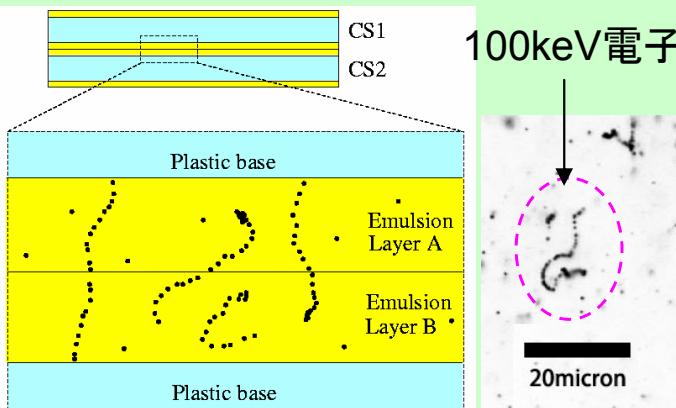


流し撮りによる読み取りの高速化

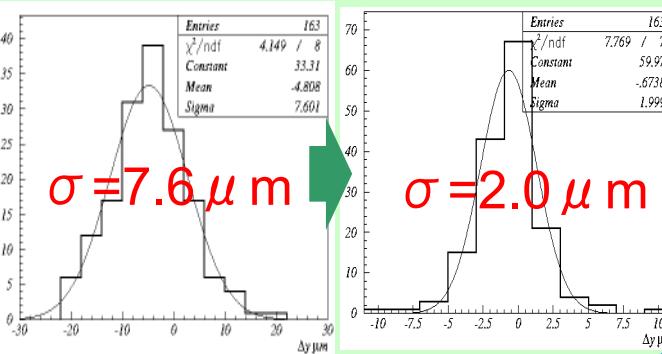


## Compton Alignment Precision alignment x 10

環境放射線によるノイズ飛跡を用いて  
フィルム間の精密アライメント

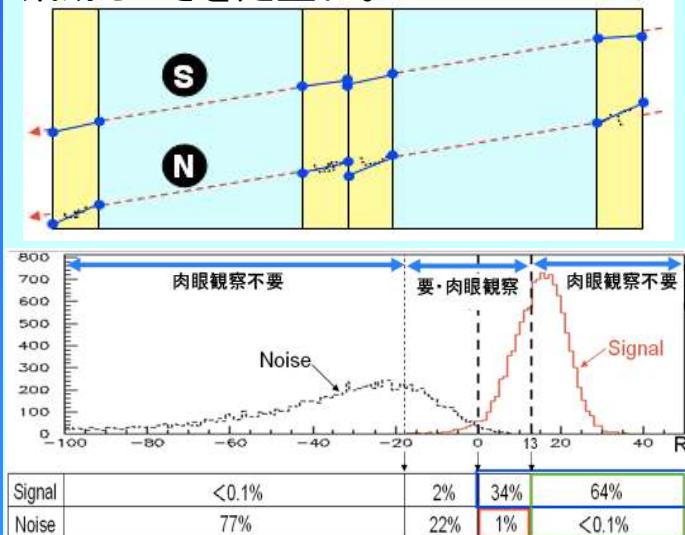


Allowance:  $30 \mu\text{m} \rightarrow 10 \mu\text{m}$

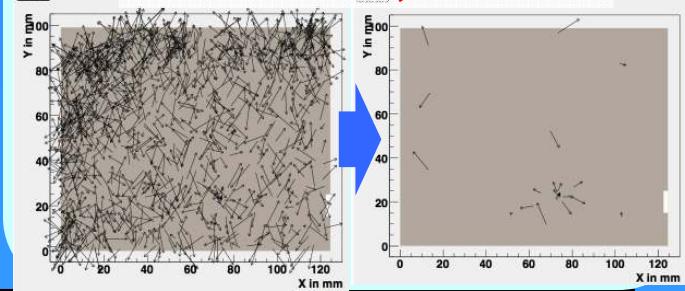


## Track Ranking High S/N discrimination x 100

直線性と濃さを総合評価し、シグナル  
飛跡らしさを定量化。

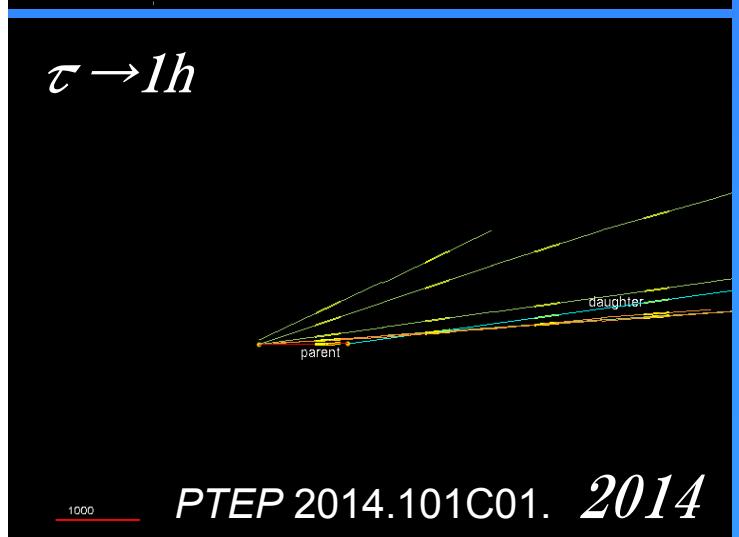
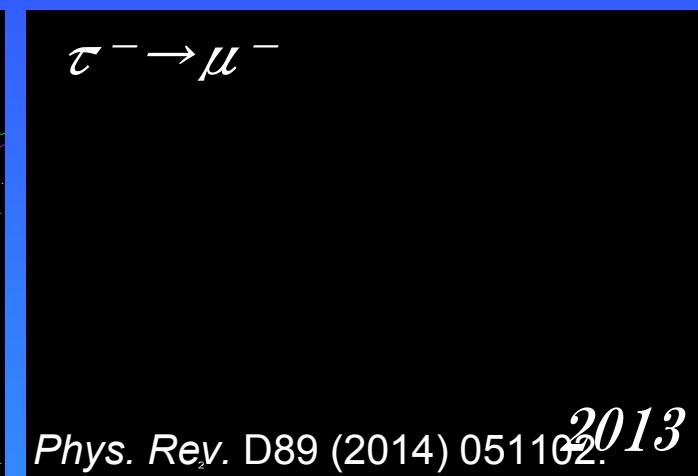
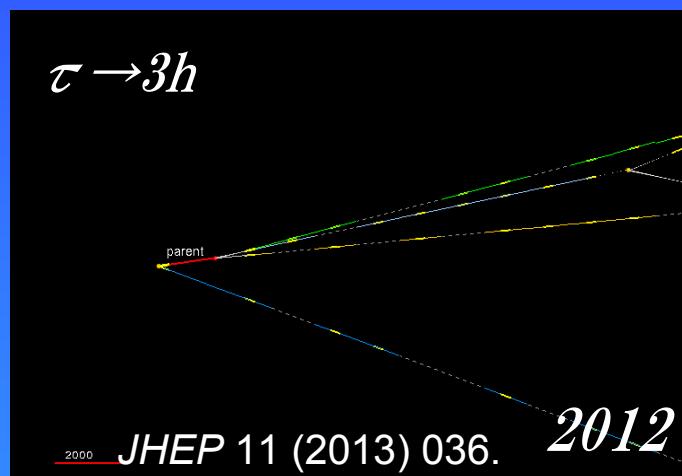
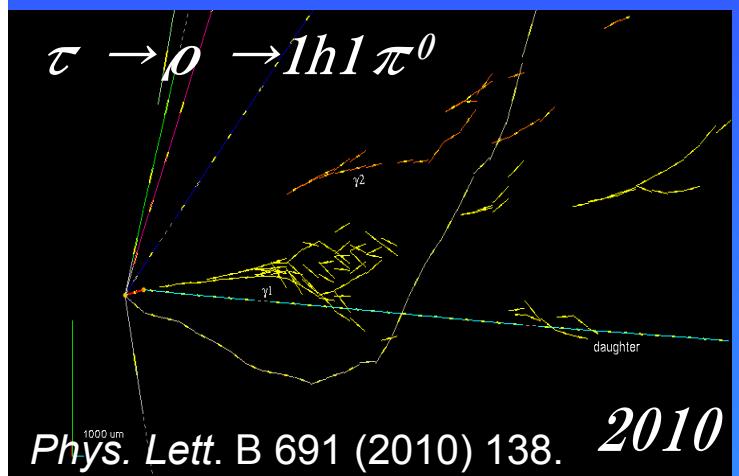
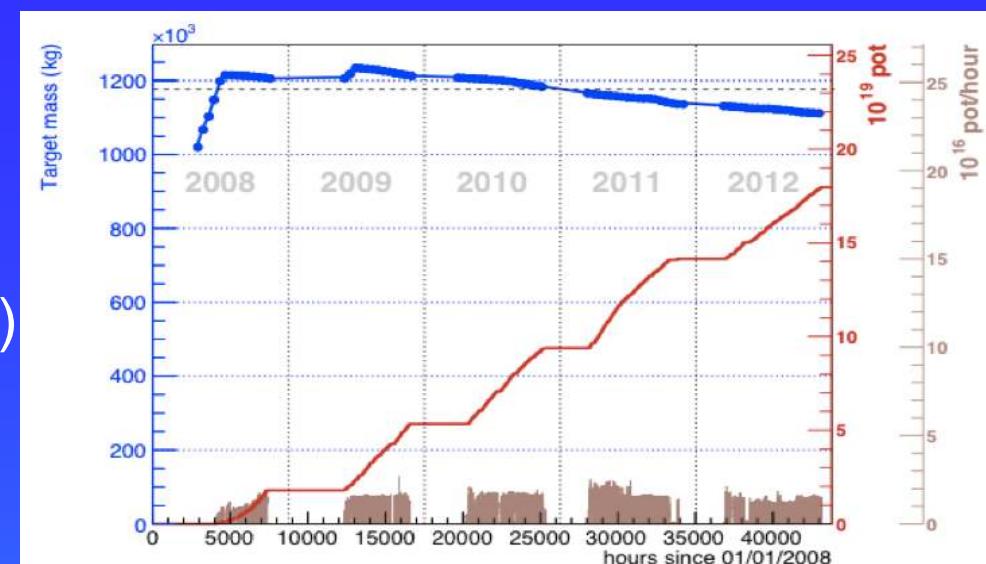


シグナル98%, ノイズ1%



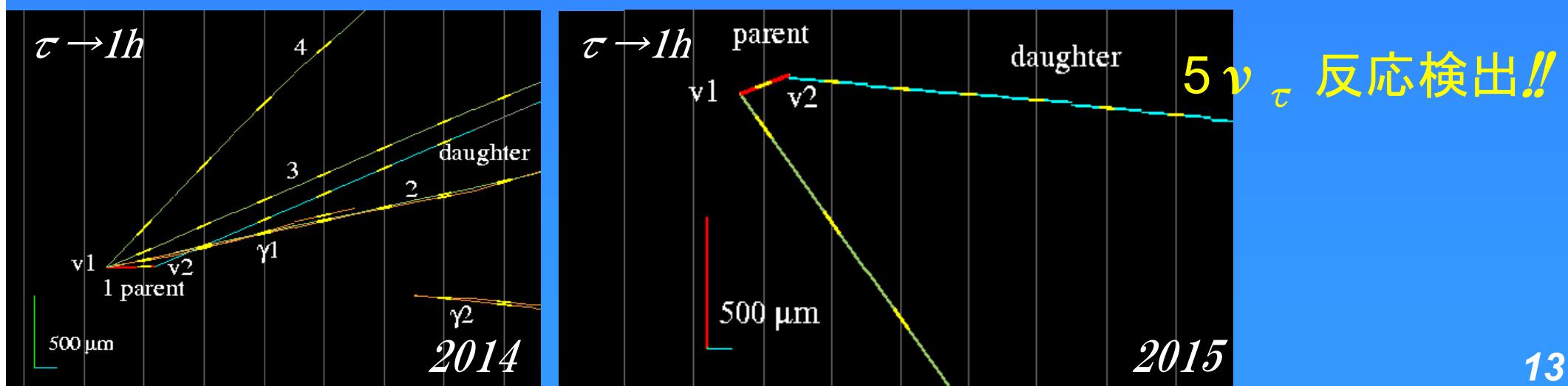
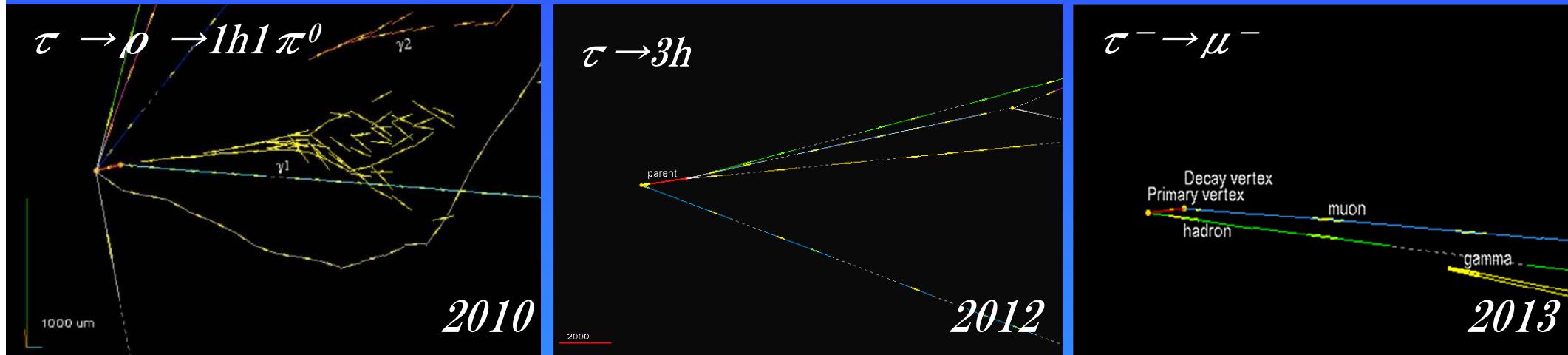
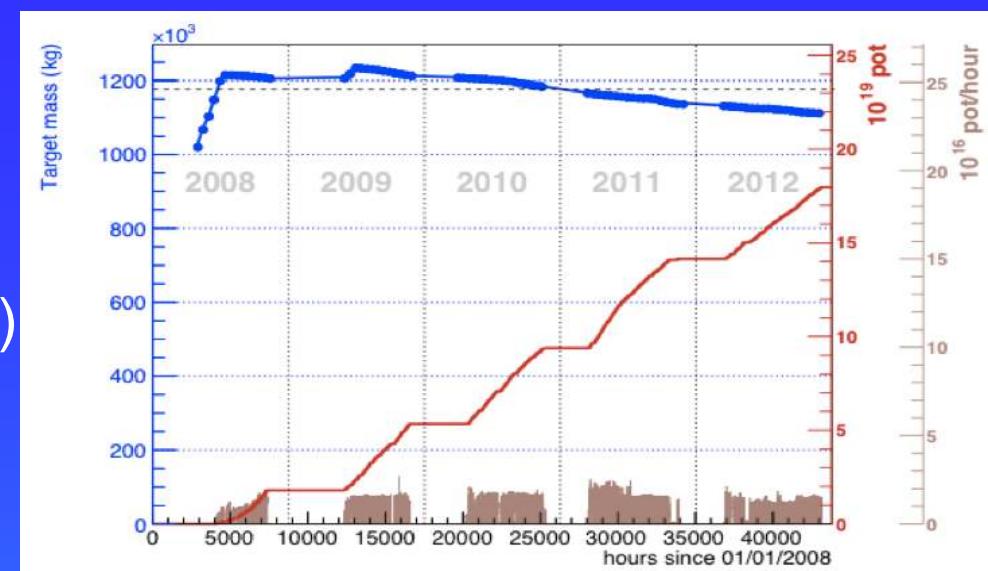
# Analysis status

- $\nu$  beam: 5 year (965 days),  
 $17.97 \times 10^{19}$  p.o.t. (80% of proposal)
- 7041 neutrino events located.  
6682 events decay searched.

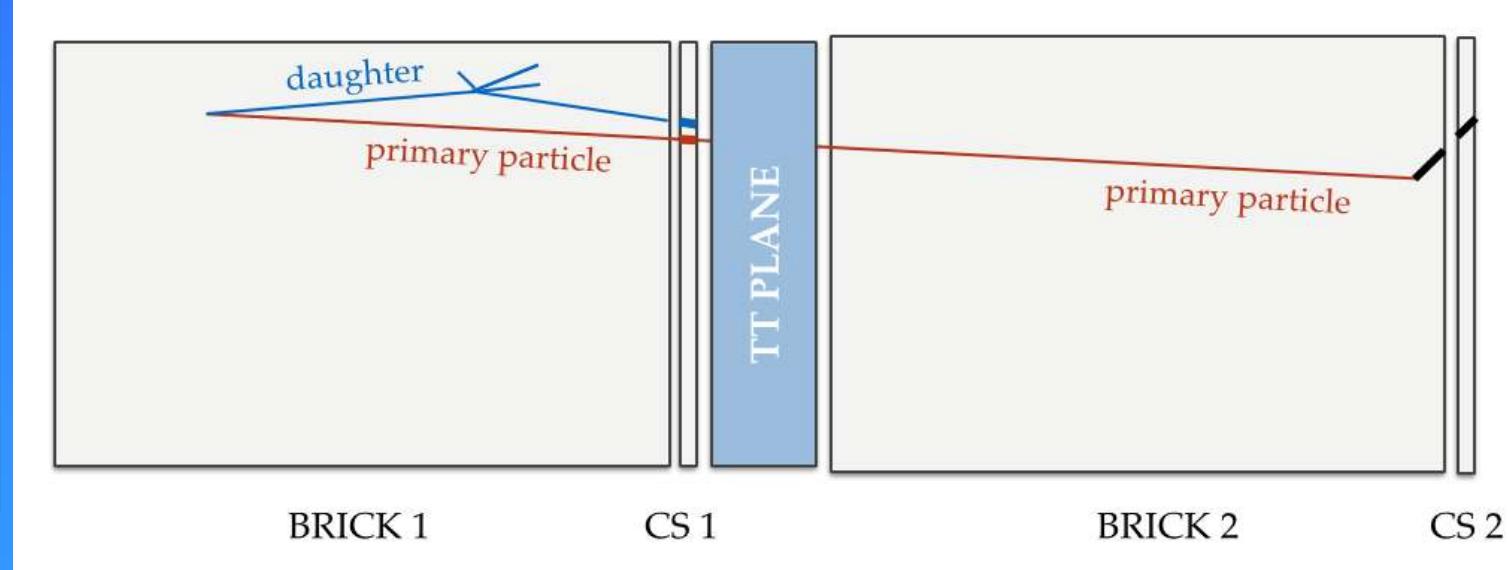
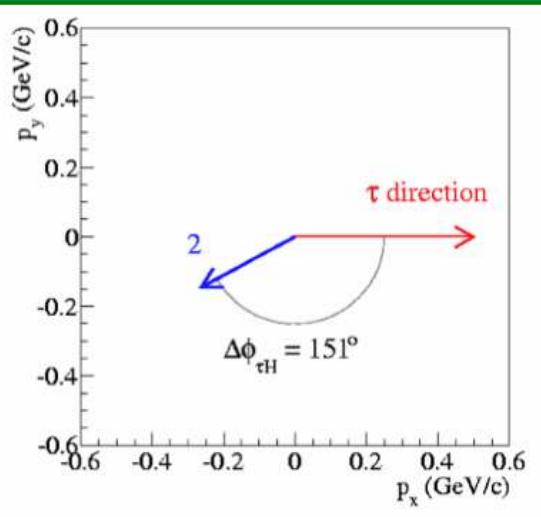
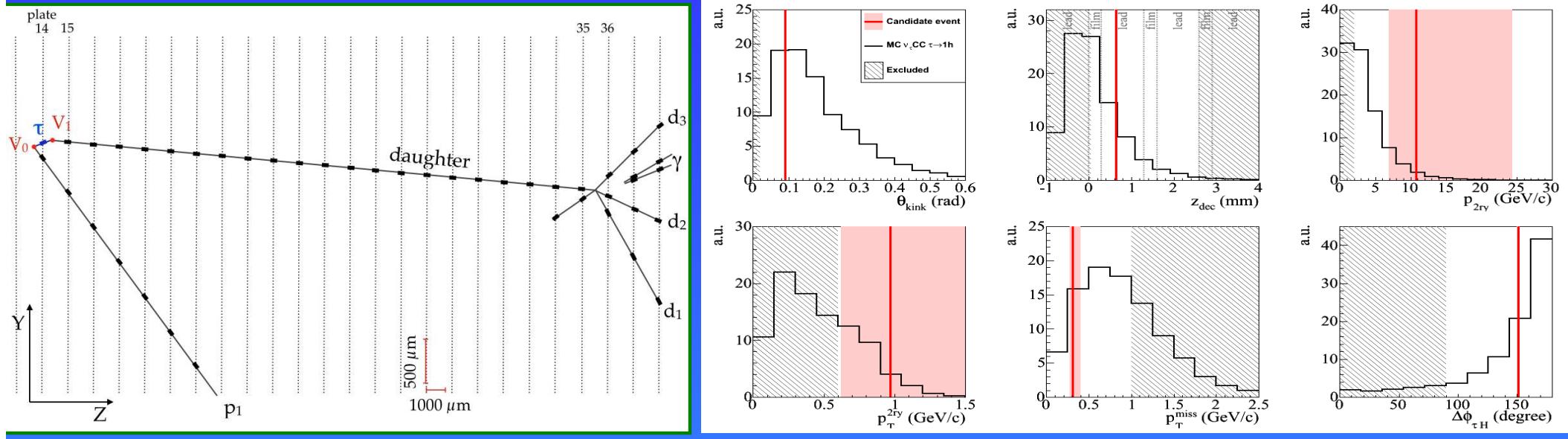


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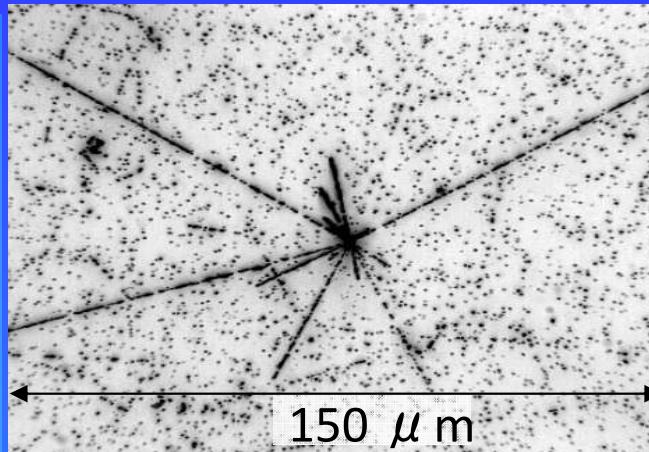
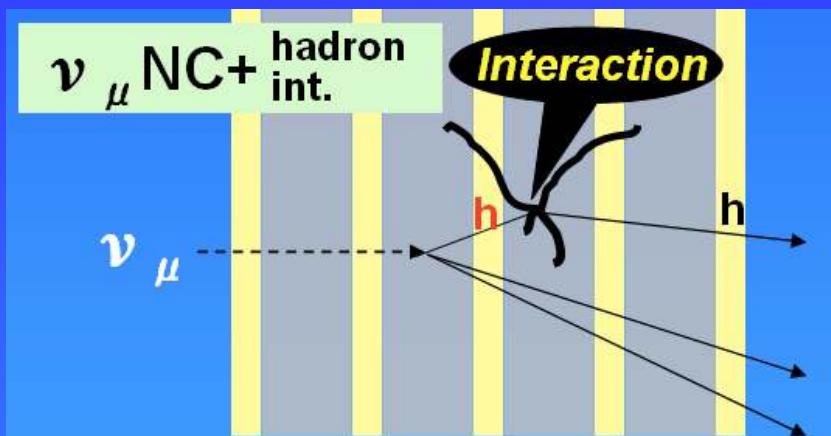
# 5<sup>th</sup> $\nu_\tau$ event



- 全てのTopology, kinematical cut をクリア

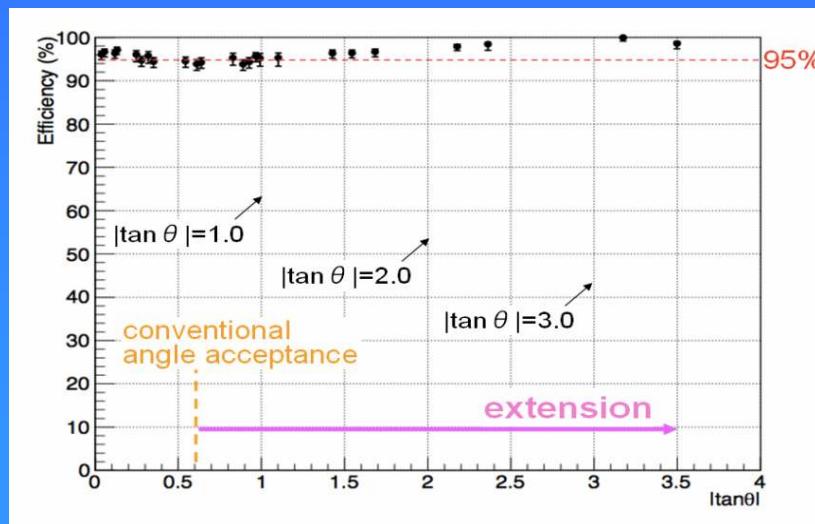
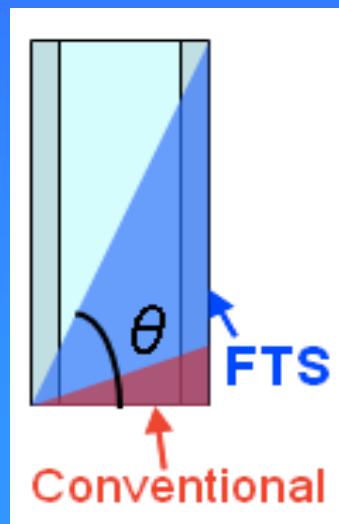
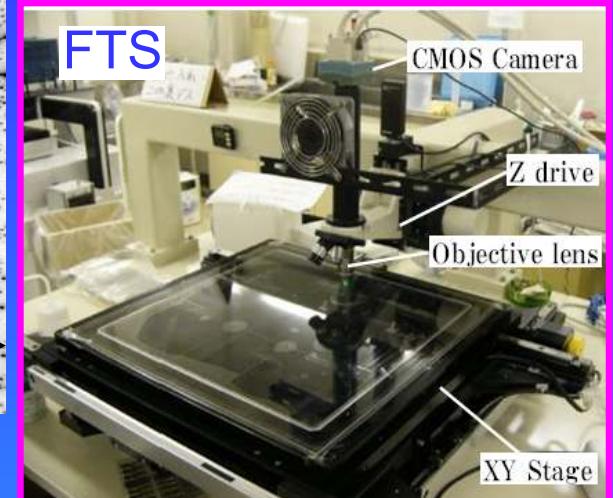
# Background reduction

- 核破砕片探索によるハドロンの2次反応バックグラウンドの低減

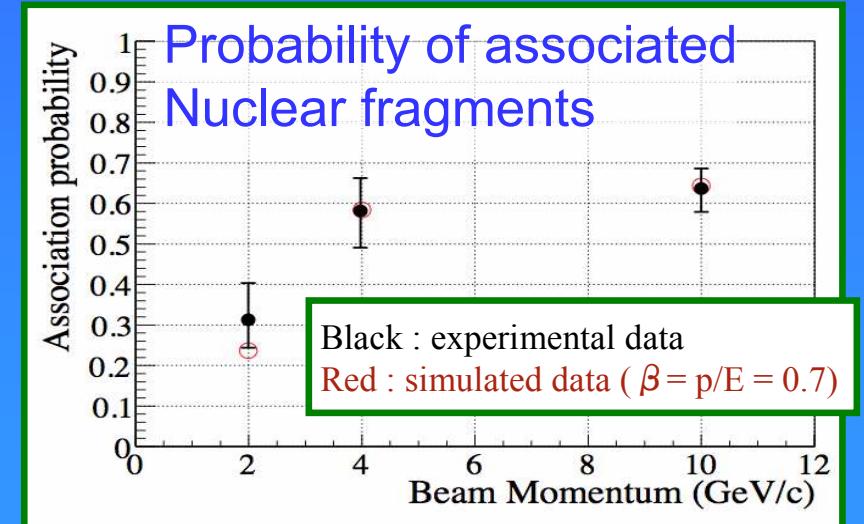


核破砕片があればハドロン2次衝突反応→無ければ崩壊事象

大角度飛跡自動認識装置の開発



T.Fukuda et al., JINST 9:P12017 (2014).



H. Ishida et al., PTEP 2014. 093C01.

大角度飛跡を自動検出できる装置を新たに開発し、実際にハドロン反応の核破砕片付随率を実験で導出。検出したタウニュートリノ反応に適用し、全てのイベントに核破砕片が付いていないことを確認した。

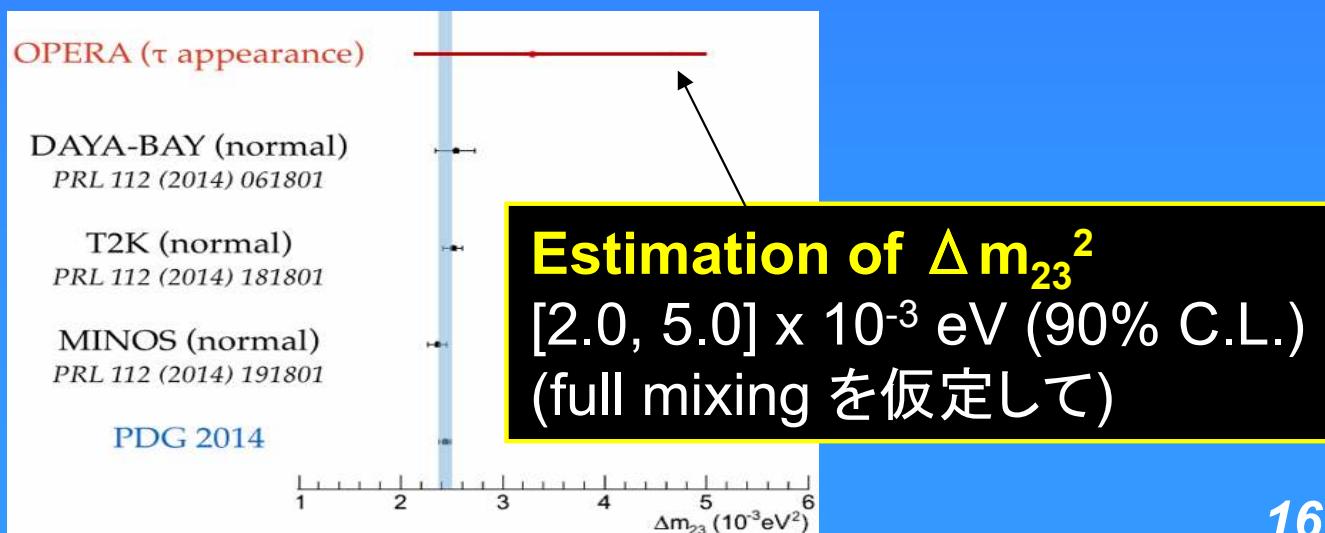
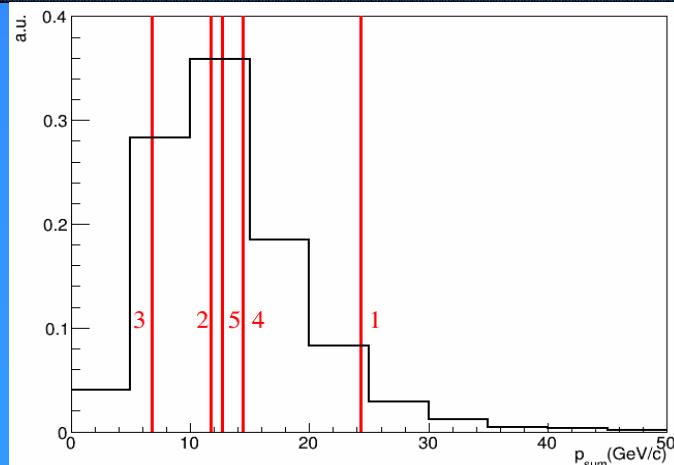
# Oscillation analysis

Expected signal and background events for the analyzed data sample

Channel	Expected background				Expected signal	Observed
	Charm	Had. re-interac.	Large $\mu$ -scat.	Total		
$\tau \rightarrow 1h$	$0.017 \pm 0.003$	$0.022 \pm 0.006$	—	$0.04 \pm 0.01$	$0.52 \pm 0.10$	3
$\tau \rightarrow 3h$	$0.17 \pm 0.03$	$0.003 \pm 0.001$	—	$0.17 \pm 0.03$	$0.73 \pm 0.14$	1
$\tau \rightarrow \mu$	$0.004 \pm 0.001$	—	$0.0002 \pm 0.0001$	$0.004 \pm 0.001$	$0.61 \pm 0.12$	1
$\tau \rightarrow e$	$0.03 \pm 0.01$	—	—	$0.03 \pm 0.01$	$0.78 \pm 0.16$	0
Total	$0.22 \pm 0.04$	$0.02 \pm 0.01$	$0.0002 \pm 0.0001$	$0.25 \pm 0.05$	$2.64 \pm 0.53$	5

- 5  $\nu_\tau$  events 検出。BG=0.25 events.  
 → バックグラウンドで説明できる確率 =  $1.1 \times 10^{-7}$   
 → **Significance = 5.1  $\sigma$ , Discovery of  $\nu_\tau$  Appearance !**

Energy distribution of  $\nu_\tau$  events



*"For the greatest benefit to mankind"*  
Alfred Nobel

2015 NOBEL PRIZE IN PHYSICS

Takaaki Kajita  
Arthur B. McDonald



## Scientific Background on the Nobel Prize in Physics 2015

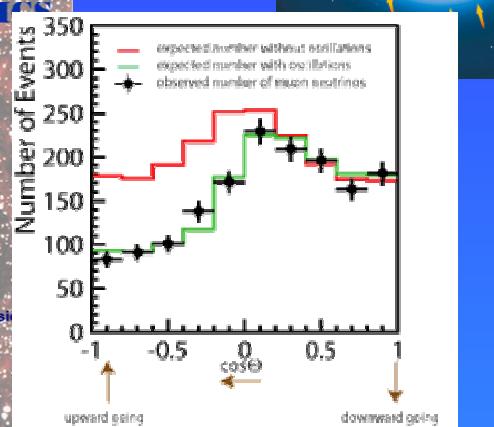
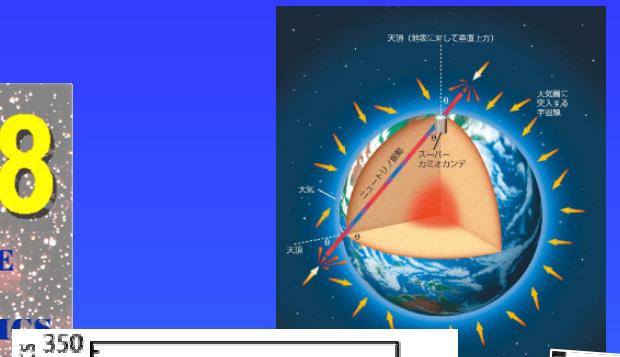
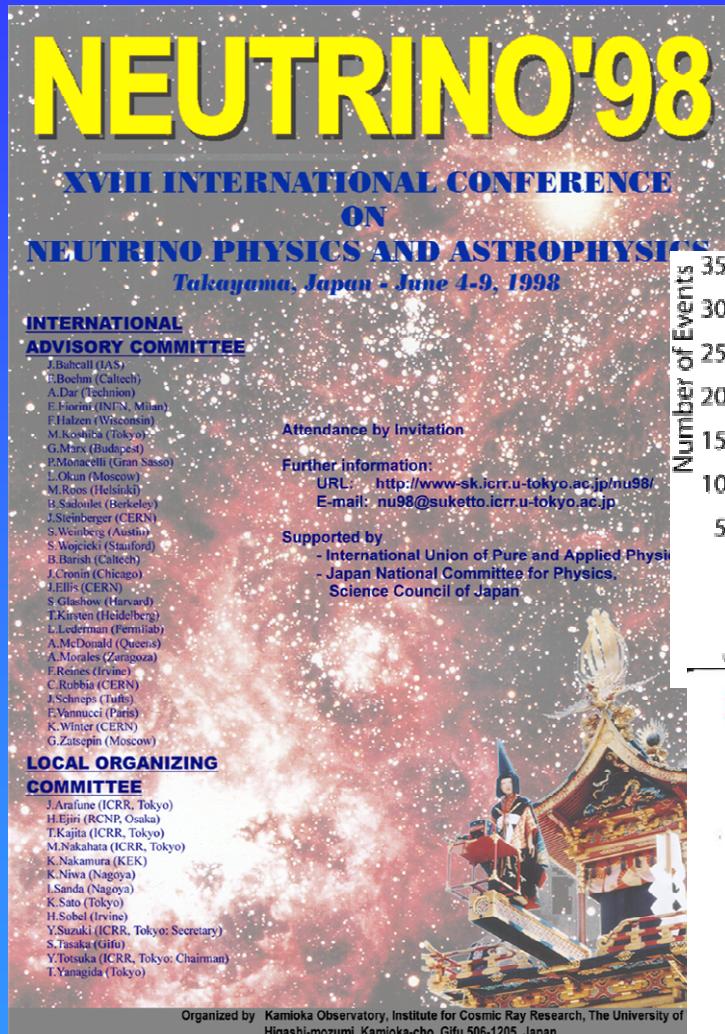
### NEUTRINO OSCILLATIONS

compiled by the Class for Physics of the Royal Swedish Academy of Sciences

n-

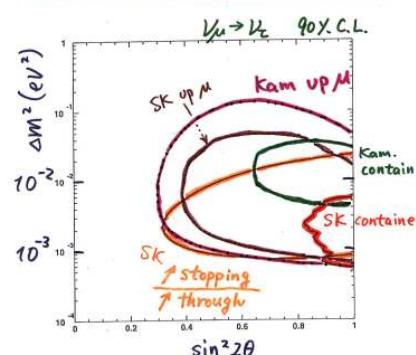
Super-Kamiokande's oscillation results were confirmed by the detectors MACRO [55] and Soudan [56], by the long-baseline accelerator experiments K2K [57], MINOS [58] and T2K [59] and more recently also by the large neutrino telescopes ANTARES [60] and IceCube [61]. Appearance of tau-neutrinos in a muon-neutrino beam has been demonstrated on an event-by-event basis by the OPERA experiment in Gran Sasso, with a neutrino beam from CERN [62].

# Looking back to XX century...



## Summary

### Evidence for $\nu_\mu$ oscillations



- $\{ \sin^2 2\theta > 0.8 \}$
- $\Delta m^2 \sim 10^{-3} \sim 10^{-2}$

(•  $\nu_\mu \rightarrow \nu_e$  or  $\nu_\mu \rightarrow \nu_s$  ?)

$\nu_\mu \leftrightarrow \nu_e$  VS  $\nu_\mu \leftrightarrow \nu_s$  solutions for atmospheric  $\nu$

O.YASUDA  
(Tokyo Metropolitan Univ.)

## 1. Introduction

\* Recent SK atmospheric  $\nu$  data  $\rightarrow$   $\left\{ \begin{array}{l} \nu \text{ oscil. w/} \\ 10^{-3} \text{ eV}^2 \lesssim \Delta m^2 \lesssim 10^{-2} \text{ eV}^2 \\ \sin^2 2\theta \simeq 1 \end{array} \right.$  isn't a [12]

## 4. Conclusions

So far both  $\nu_\mu \leftrightarrow \nu_e$  &  $\nu_\mu \leftrightarrow \nu_s$  solutions provide a good fit to atmospheric neutrino anomaly

for  $\left\{ \begin{array}{l} 10^{-3} \text{ eV}^2 \lesssim |\Delta m^2| \lesssim 10^{-2} \text{ eV}^2 \\ \sin^2 2\theta \sim 1 \end{array} \right.$ .

To be more conclusive  
we need more statistics  
or we have to look for  
appearance of  $\nu_e$  in long  
baseline experiments.

done

Next step ... !!

# **Status of J-PARC T60**

# J-PARC T60

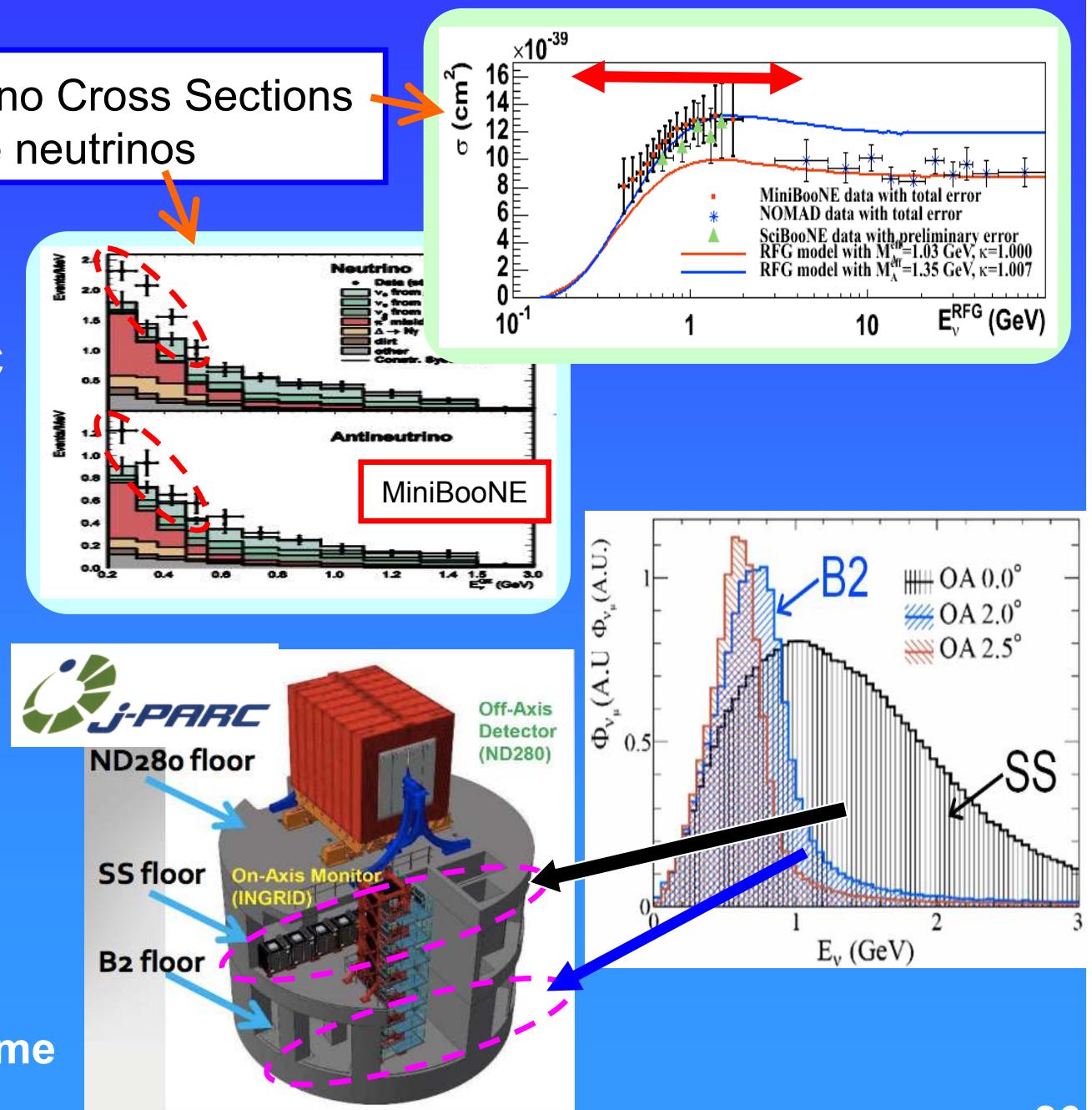
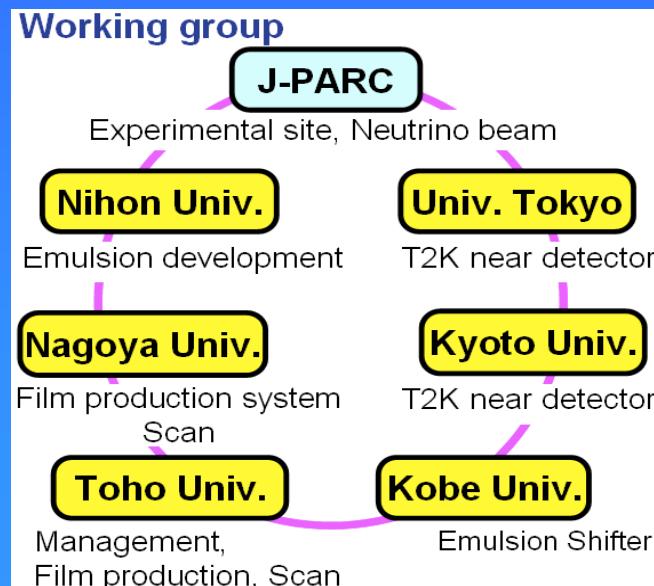
PI : T. Fukuda (Toho Univ.)  
 → tsutomu.fukuda@ph.sci.toho-u.ac.jp

## Future Physics goals

- Neutrino Cross Sections
- Sterile neutrinos

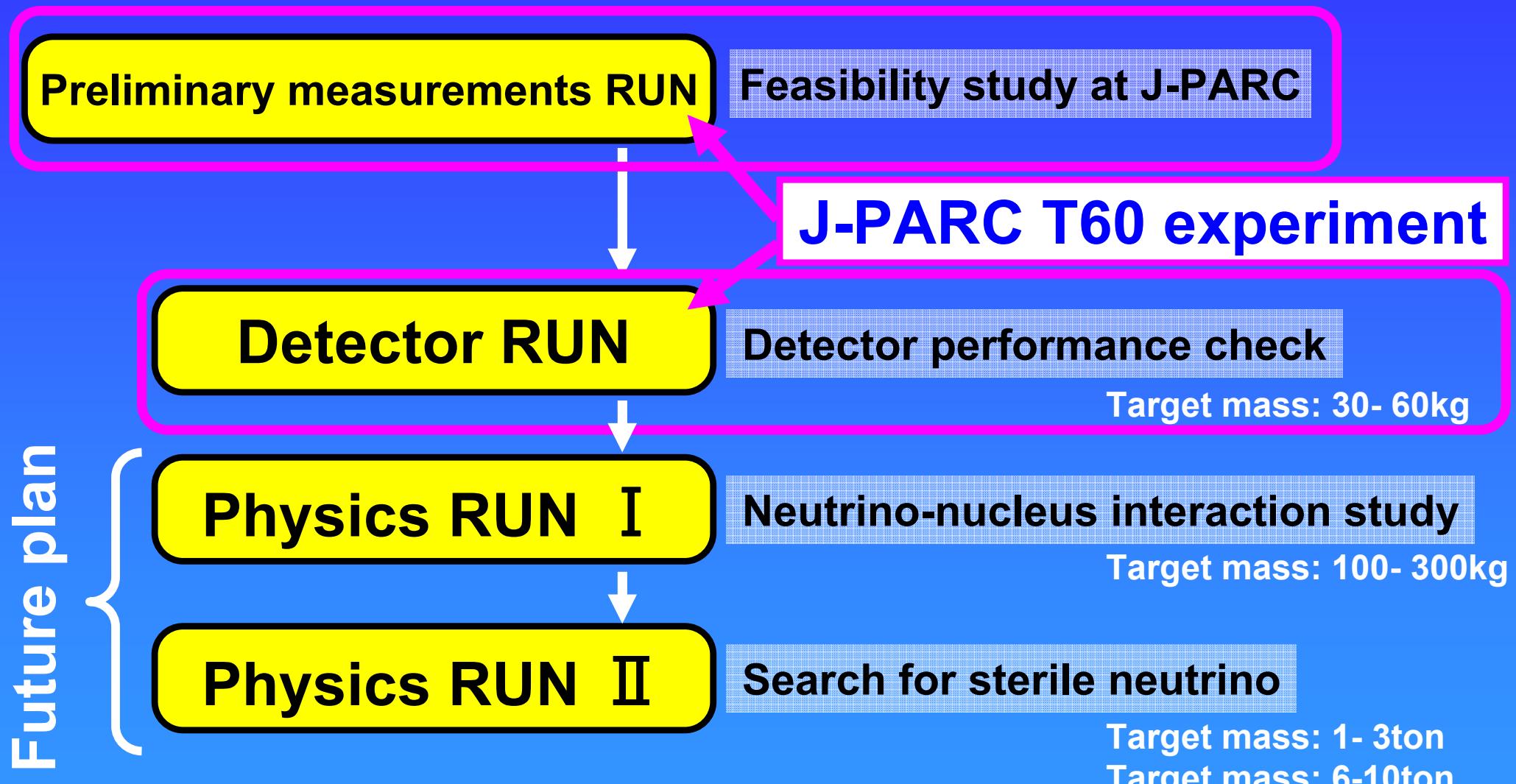
New experimental program to study low energy neutrino int. with nuclear emulsion at J-PARC was proposed.

T60 is a first step of this series.



A collaborative project with some member of OPERA and T2K

# *Roadmap*



- The aim of T60 is a **feasibility study** and **detector performance check** to make a future plan.
- We will expand the scale of detector gradually, step by step.

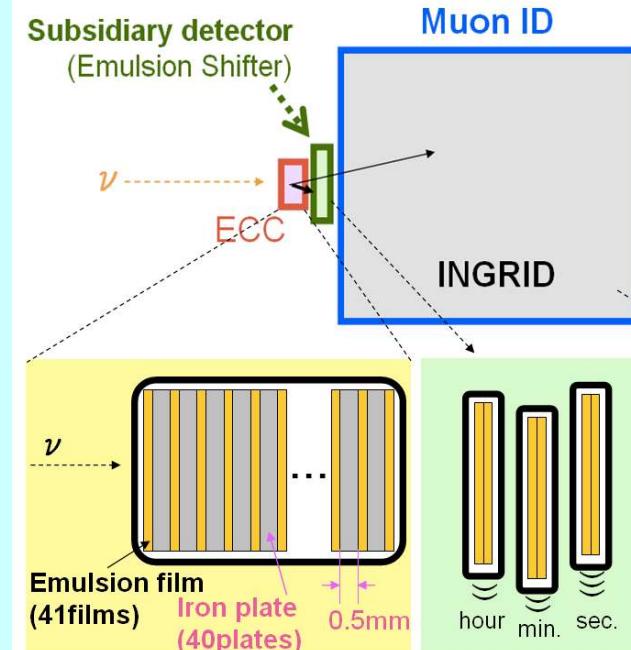
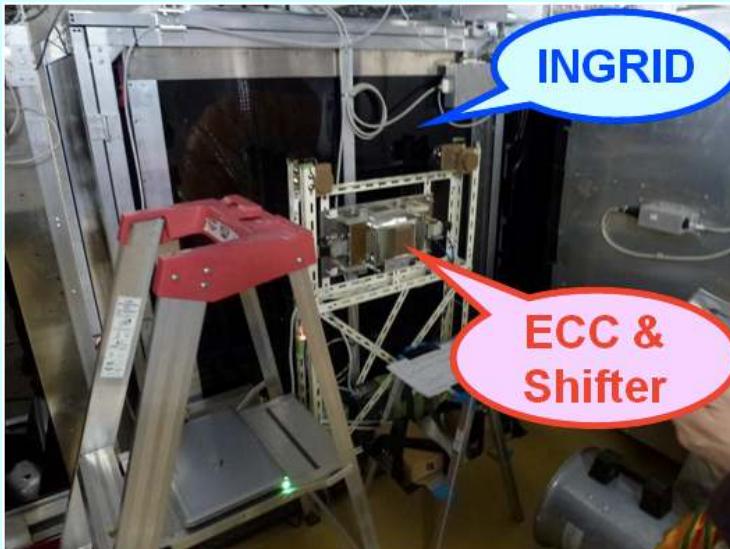
# Status of T60

Feasibility study: 2kg Iron target ECC

## Emulsion film production by ourselves

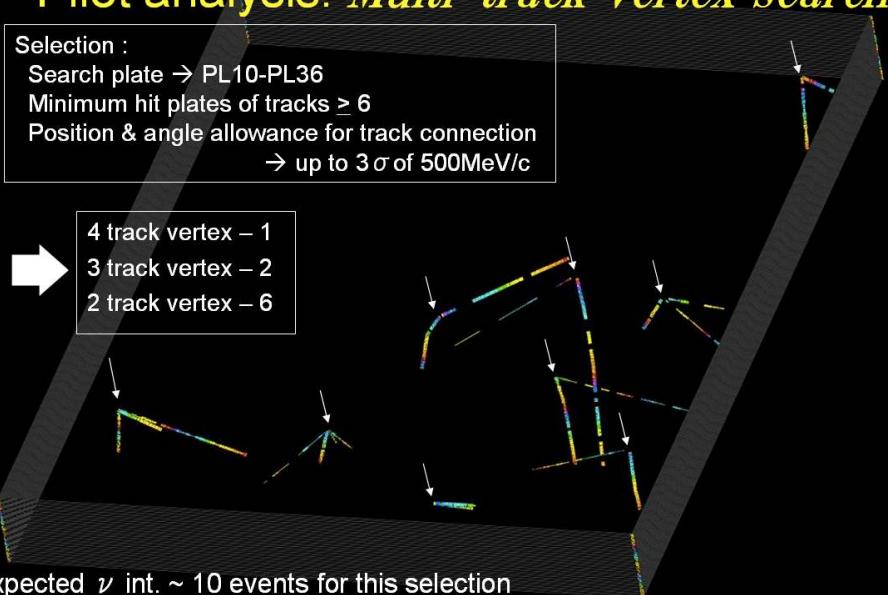


## Installation @J-PARC

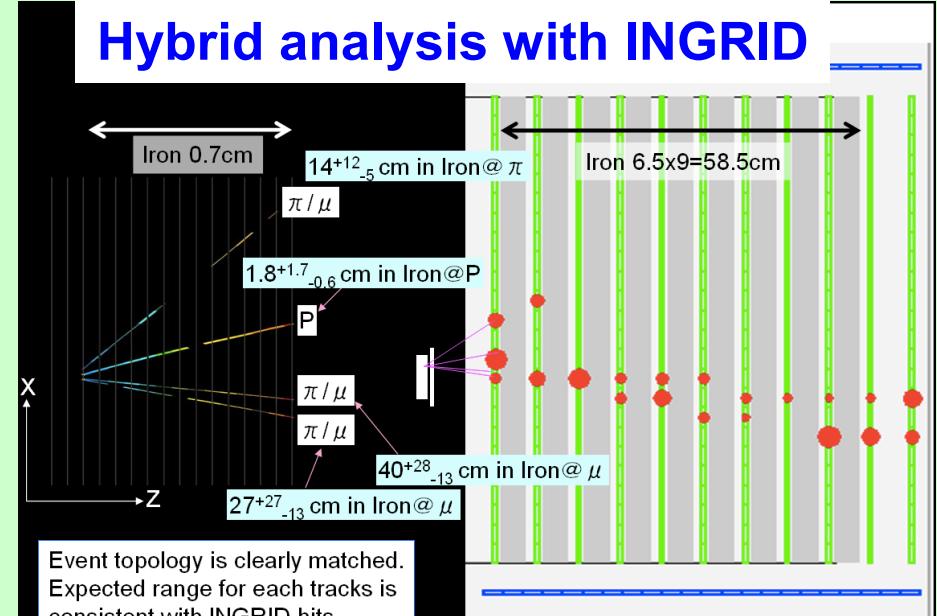


## Pilot analysis: Multi-track vertex search

Selection :  
Search plate → PL10-PL36  
Minimum hit plates of tracks  $\geq 6$   
Position & angle allowance for track connection  
 $\rightarrow$  up to  $3\sigma$  of 500MeV/c



## Hybrid analysis with INGRID



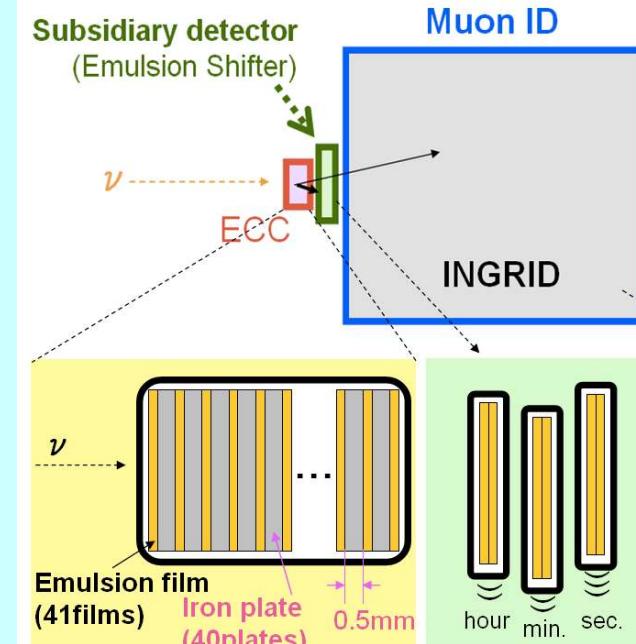
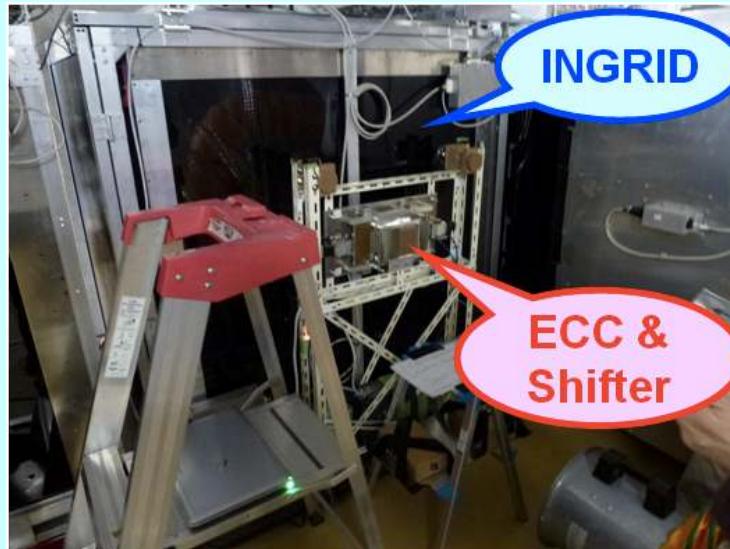
# Status of T60

Feasibility study: 2kg Iron target ECC

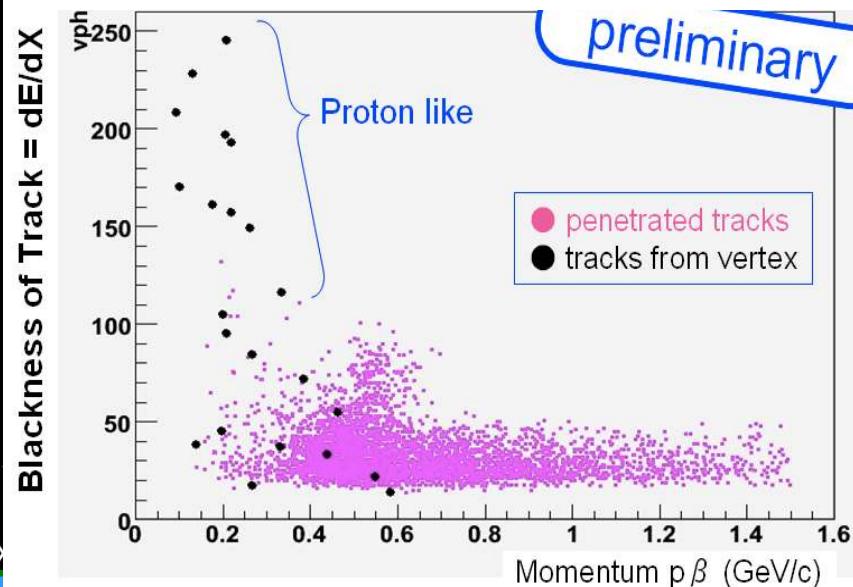
## Emulsion film production by ourselves



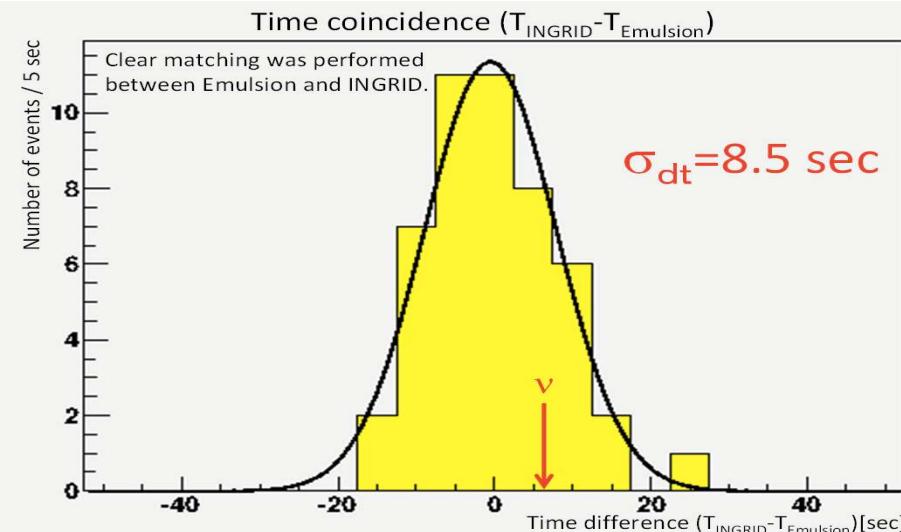
## Installation @J-PARC



## Proton identification



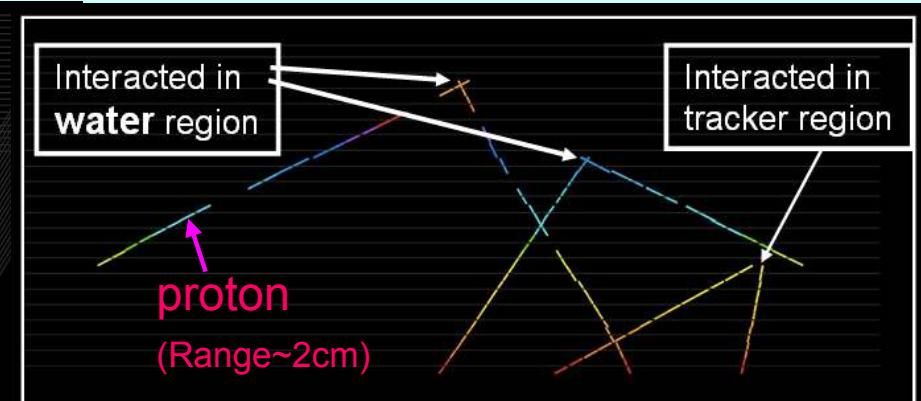
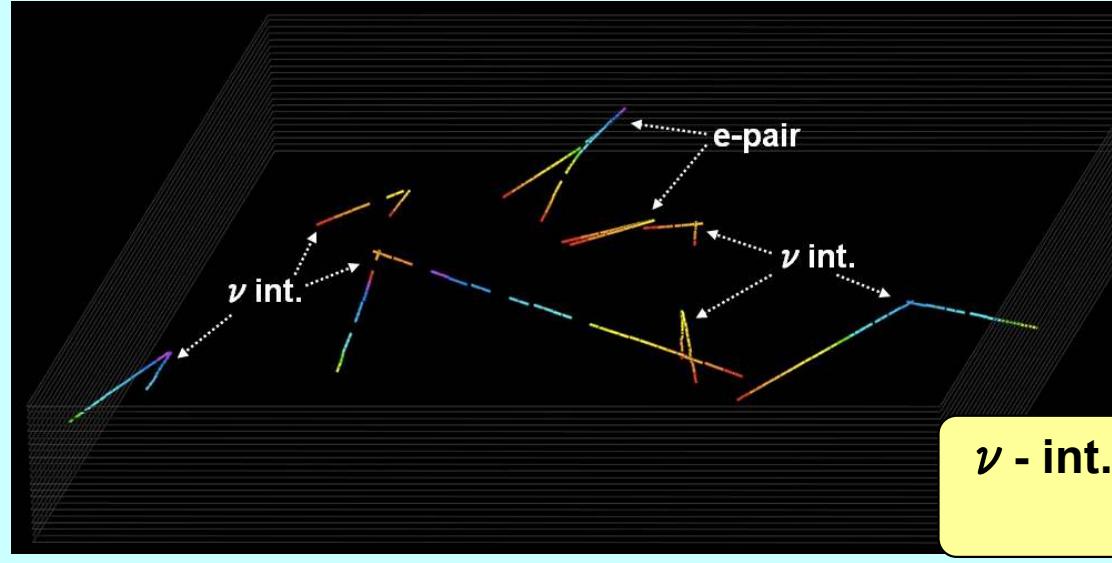
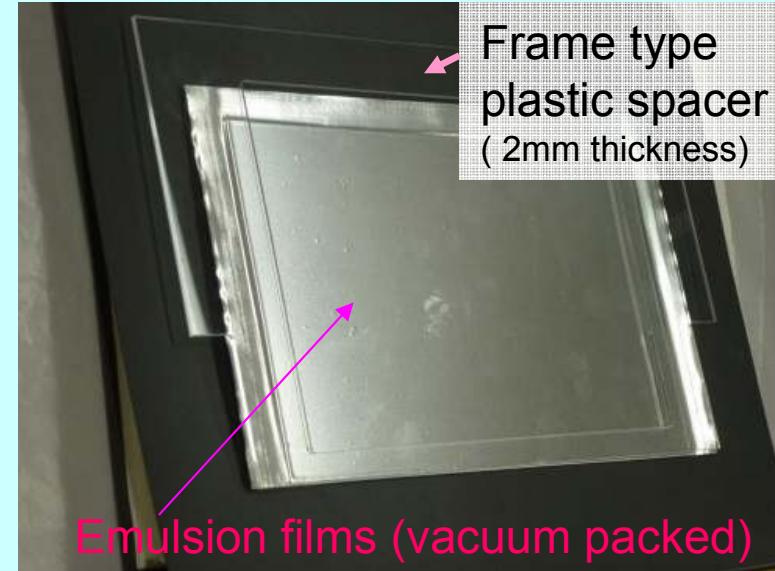
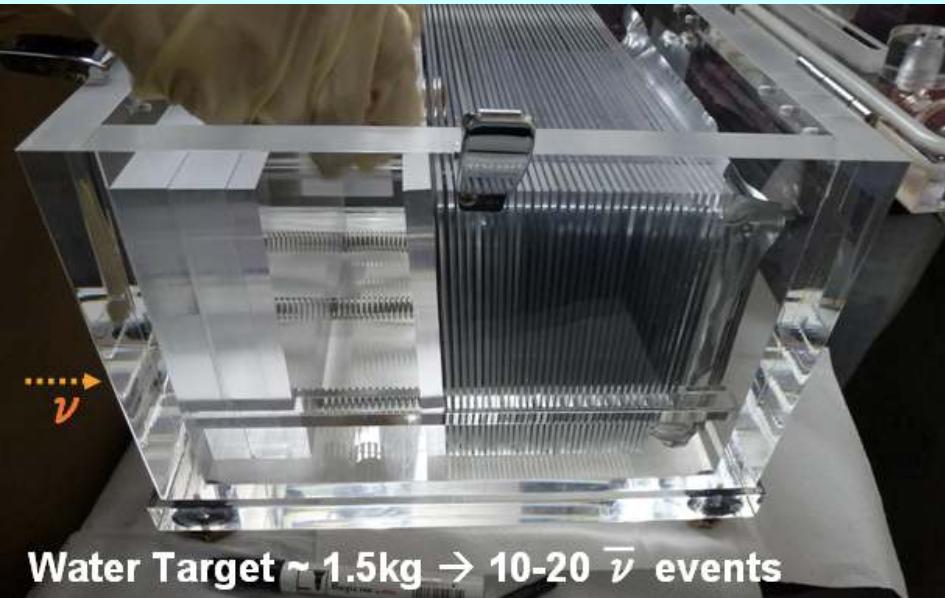
## Time resolution



# Status of T60

Feasibility study: 1.5kg Water target ECC

## Water target emulsion chamber

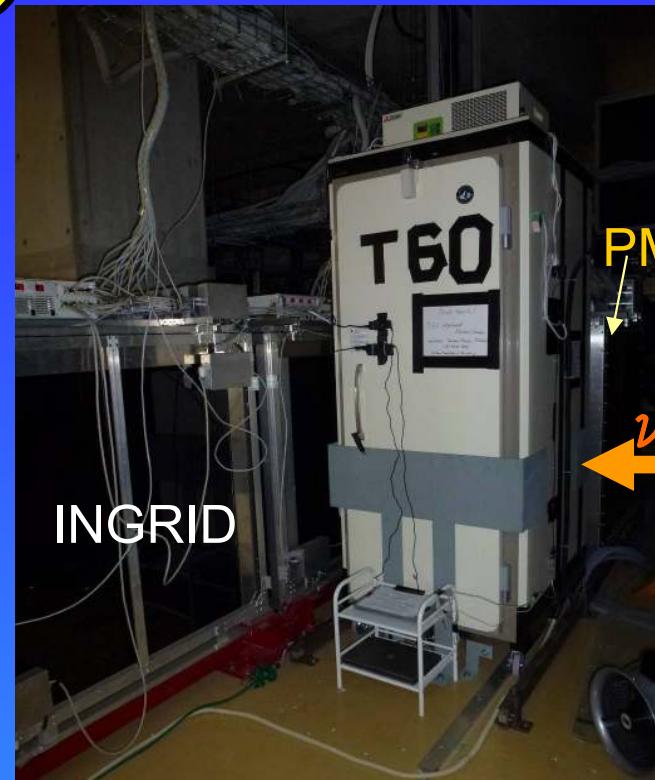
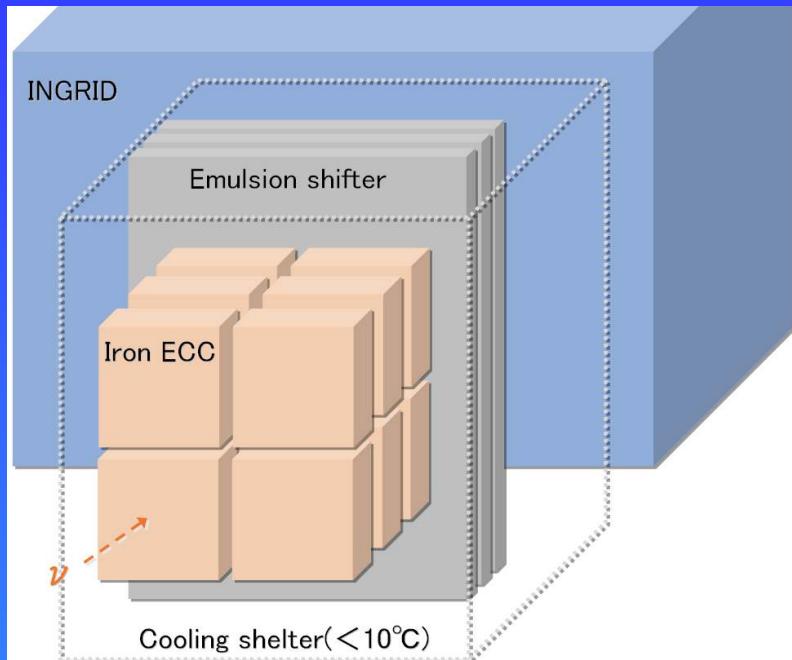


$\nu$  - int. → 6 events { Water region : 2 events  
Tracker region : 4 events

First observation of  $\nu$  - Water interactions with Emulsion Detector

# Status of T60

Detector Run: 60kg Iron target ECC



- $\bar{\nu}$  exposure : 2016 @SS floor  
end of Jan. → beam end
- Iron target (total~60kg :  $500 \mu\text{m}$  seg.)
- High statistics (4-6k  $\bar{\nu}_\mu$  events)
- $\nu_e$  detection (30-40  $\bar{\nu}_e$  CC events)

→ Data – MC comparison with high statistics to check the performance.



# *Summary*

## OPERA

- OPERA successfully collected data from 2008 to 2012. A total number of  $17.97 \times 10^{19}$  p.o.t. integrated ( $\sim 80\%$  of the nominal value).
- **5  $\nu_\tau$  candidate events** were found with 2.6 signal and 0.25 background events expected in the analyzed sample.
- Significance of the observation is **5.1  $\sigma$**   
→ **Discovery of  $\nu_\tau$  appearance** in the CNGS beam.

## J-PARC

- We are planning neutrino experiments **at J-PARC** to study low energy neutrino - nucleus interactions with nuclear emulsion.
- First of all, we are carrying out a test experiment at J-PARC (**T60**) to check the feasibility and detector performance.
- We demonstrated **neutrino event analysis, hybrid analysis** with T2K near detector and detection of  **$\nu$ -water interactions**.
- We will check the detector performance with high stat. and establish a detailed plan for physics run.

# *Hadron Production measurements*

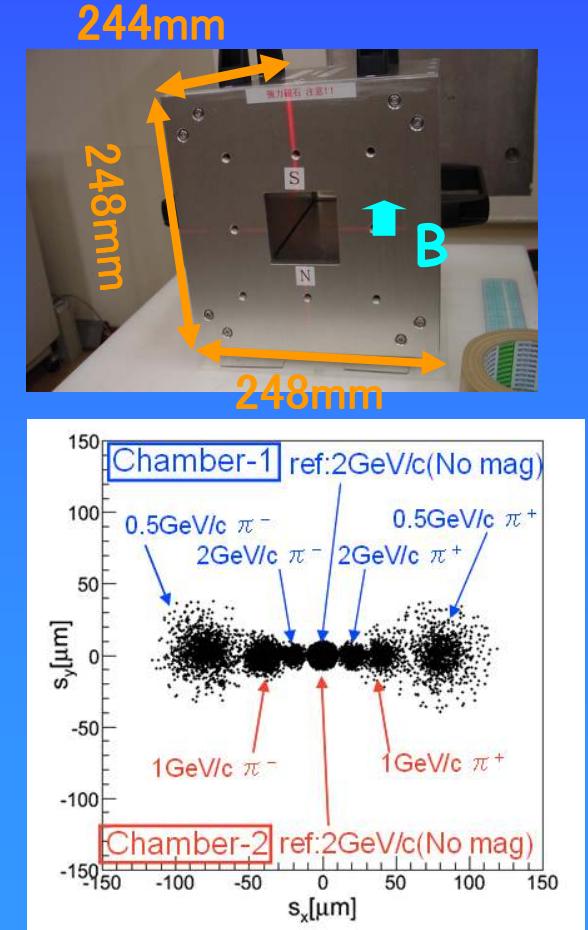
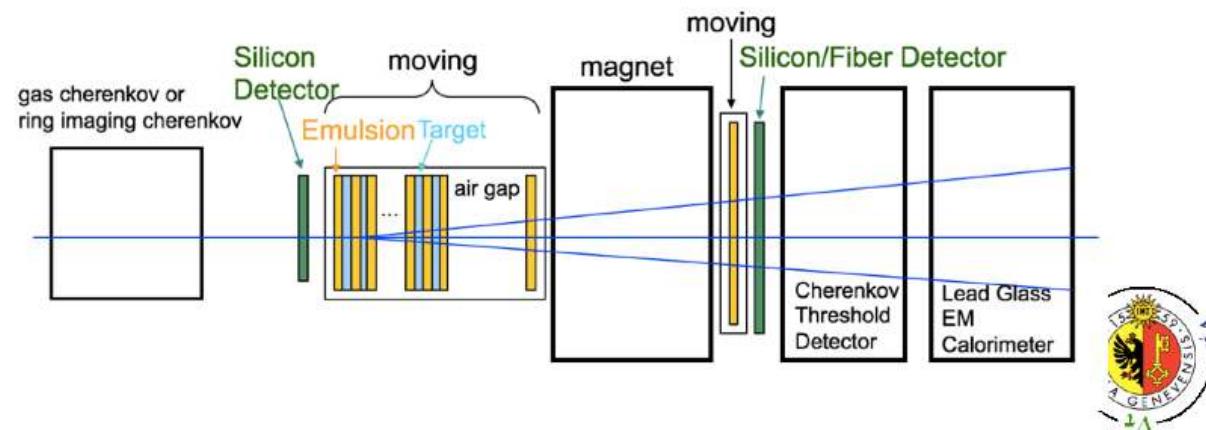
- To reduce uncertainty of neutrino flux (accelerator  $\nu$ , atmospheric  $\nu$ )
- Several – 10 GeV/c hadron interactions is studied with nuclear emulsion.
- First, target is Carbon. Then Iron, Aluminum will be performed.

## Hybrid Emulsion Detector

Detector that uses emulsion film tracking is being developed by ICRR, Kavli IPMU, Kyoto U., Toho U. + KEK, Nagoya U., Kobe U., Nihon U.

- Minimize material between tracker and target (large systematic effect for HARP)
- Compact size detector can be moved between different beam lines
- Detailed measurements of interaction topologies

Emulsion detector tracks connected to upstream and downstream particle ID detectors by silicon strip or pixel detectors



- We plan to expose hadron beam at Fermilab/CERN in the near future.