

# **XMASS experiment**

## **Current status**

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# What is XMASS ?

XMASS is a multi purpose **low background** experiment **using Liq.Xe.**

## XMASS

**Xenon MASSive** detector for Solar neutrino ( $pp/\text{}^7\text{Be}$ )

· **Low energy solar neutrinos** Plenty information for osc. parameters

**Xenon detector for Weakly Interacting MASSive Particles**  
(Dark Matter search)

· **Dark Matter** Non Baryonic Cold Dark Matter Candidate

**Neutralino**

Leading Candidate

Direct Detection of WIMPs(Neutralinos)

↳ Very important for Astrophysics and Particle physics .

**Xenon neutrino MASS** detector (double beta decay)

·  $0\nu\beta\beta$  ,  $2\nu\beta\beta$  decay

# Why liq. Xe?

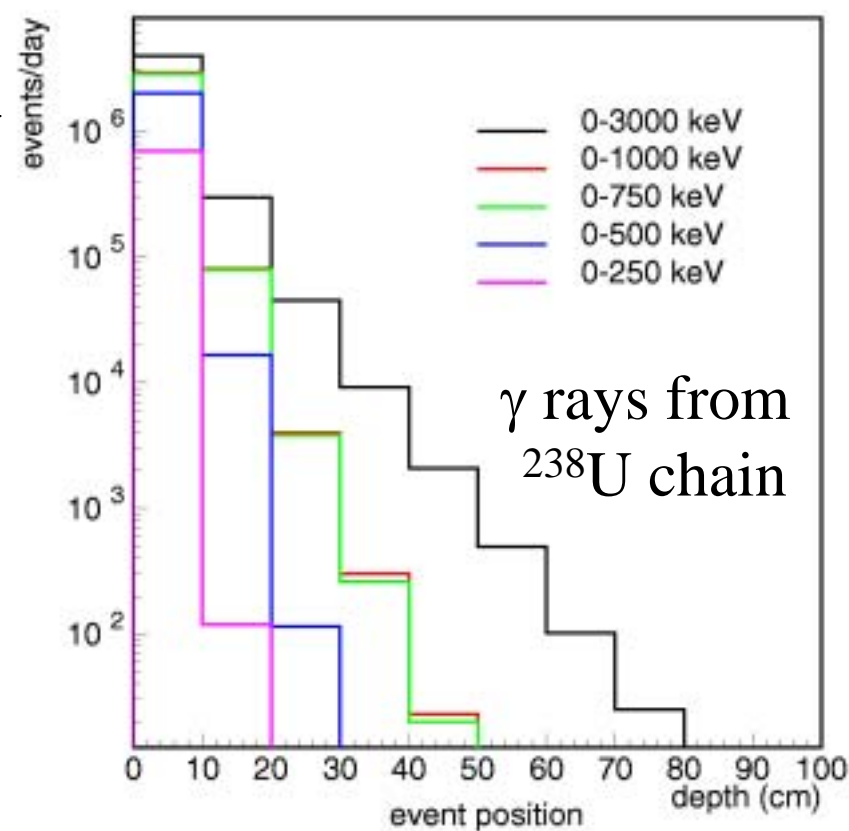
- Large photon yield (**~42000photons/MeV**)
- Compact detector size (**~3g/cm<sup>3</sup>** 10ton=1.5m cubic)
- Purification (distillation)
- No long life isotope
- Scintillation wavelength (**175nm**, detect directly by PMT)
- Relative high temperature (**~165K**)

- **Self-shield** (large **Z=54**)

**Several orders of magnitude reduction** can be expected for energy less than 250keV at 20cm

- **easy isotope separation**

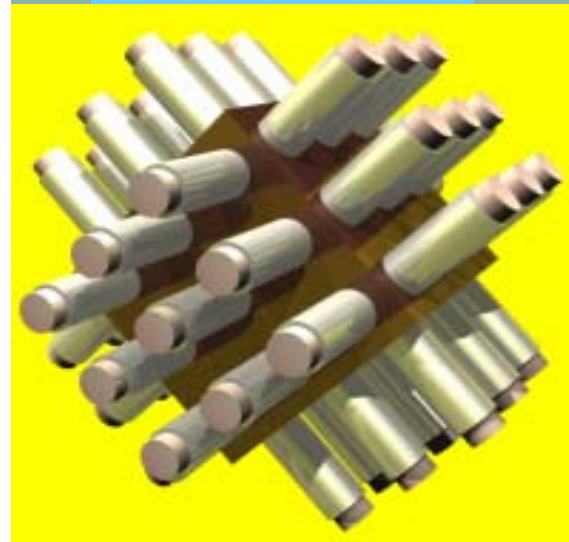
<sup>136</sup>Xe : double beta decay  
odd-Xe : WIMPs spin Dep. interaction  
even-Xe : WIMPs spin InDep. interaction



# Experimental strategy

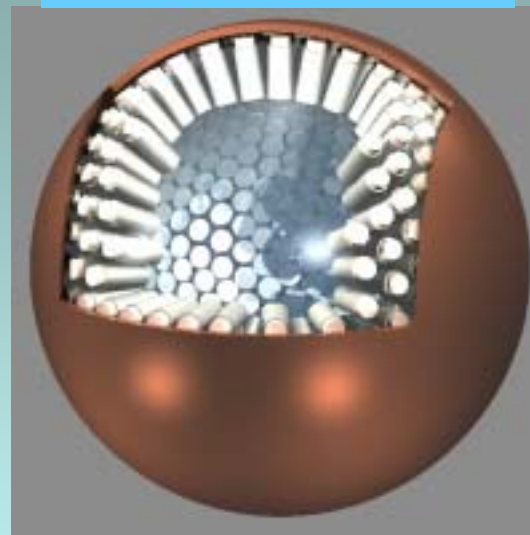
As a first step of XMASS, Dark matter search is main purpose.

100kg Detector  
Mainly R&D



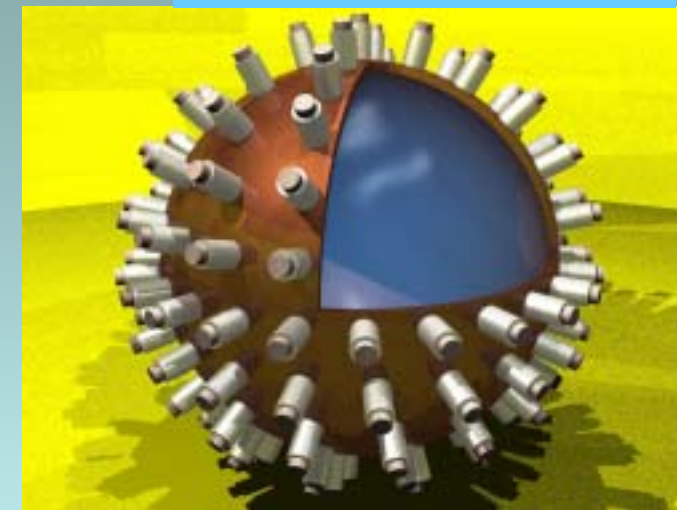
Now started!!

800kg Detector  
Dark matter search



Next planning!

~10 ton Detector  
Solar  $\nu$  / Dark matter



Performance demo., R&D

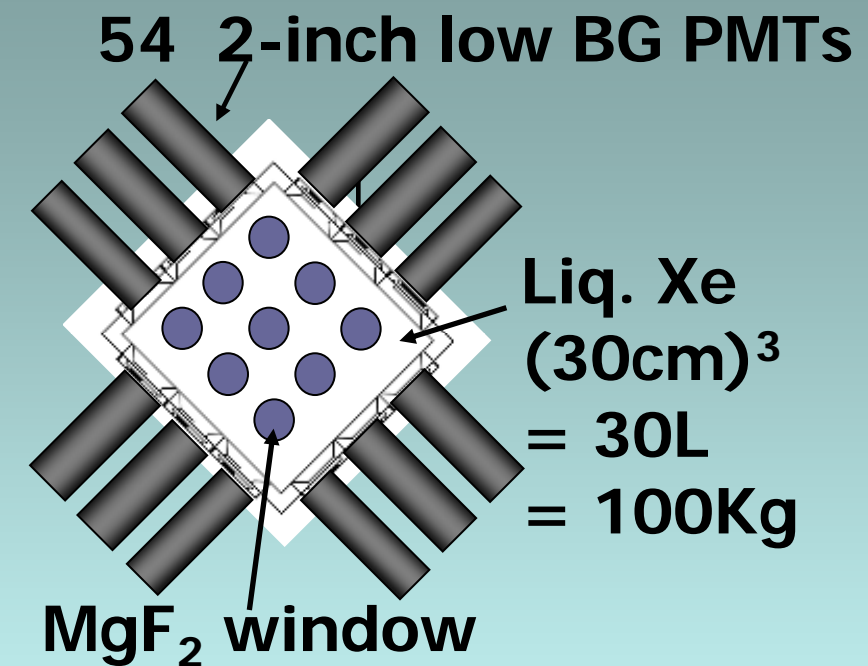
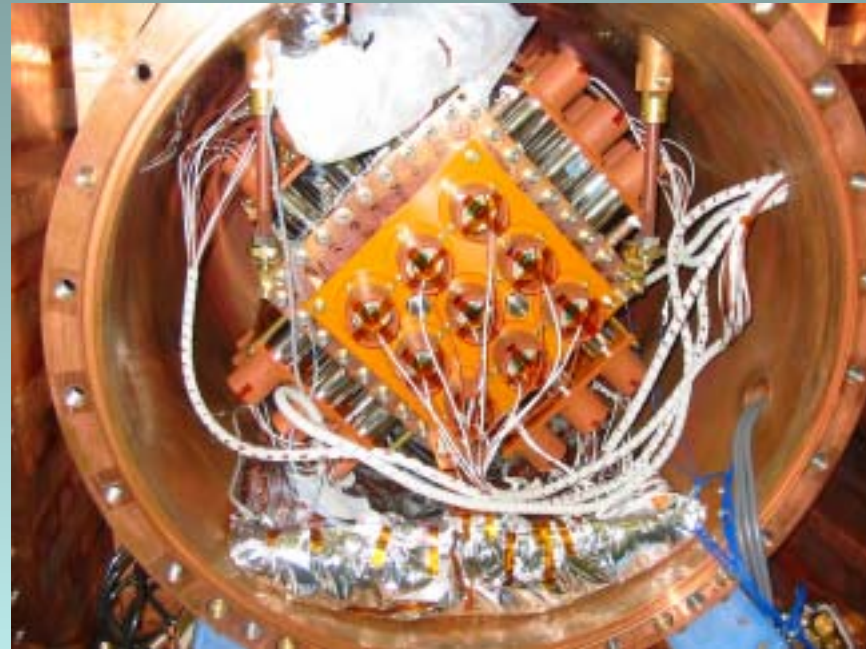
LowE solar neutrino detection

Dark matter search

R&D of the double beta dedicated detector

**100kg Detector  
current  
status/analysis**

# 1. Introduction of 100kg detector



## Menu of R&D

- Low background setup
- Vertex / energy reconstruction
- Demonstration of self-shielding
- Purification system
- attenuation length
- neutron B.G. study etc.

## 2. Low background setup

### · selection of material

Inner Vacuum Chamber  
Outer Vacuum Chamber

} made of OFC

PMT



- PMT base : glass PCB      PTFE PCB  
 $^{238}\text{U}$        $\sim 1/100$  ( $\sim 10^{-3}\text{Bq/PMT}$ )
- Glass tube      metal tube

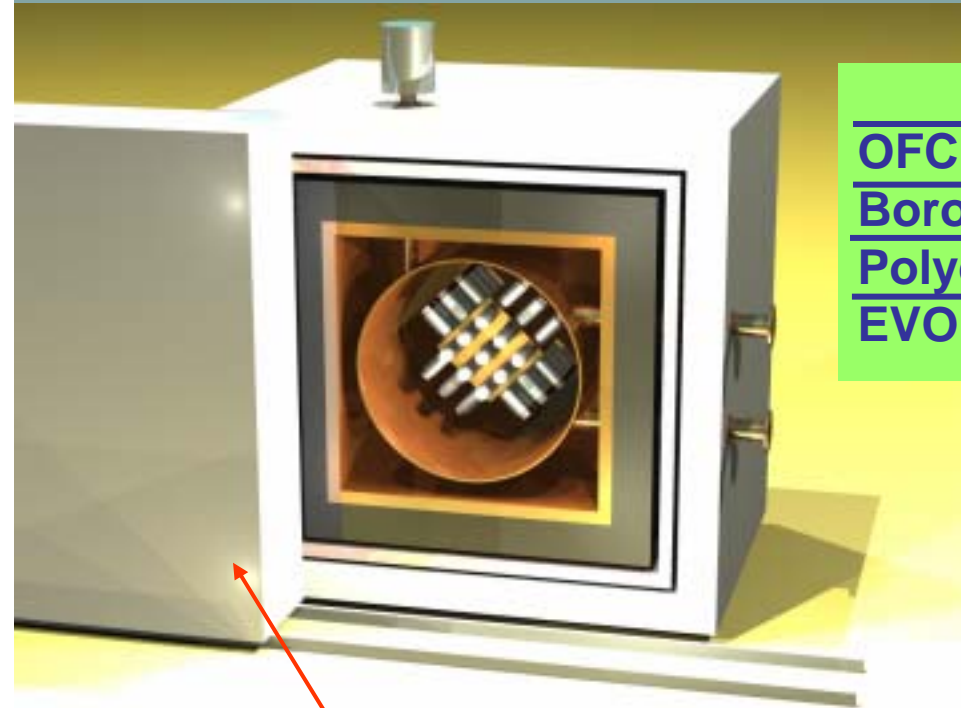
	U(Bq/PMT)	Th(Bq/PMT)	K(Bq/PMT)
R8778	$2.0 \times 10^{-2}$	$7.0 \times 10^{-3}$	$1.4 \times 10^{-1}$
ZK0667	$5.0 \times 10^{-1}$	$1.2 \times 10^{-2}$	$6.1 \times 10^{-1}$

Low background PMT !!

Other materials are also low radioactivity.

## 2. Low background setup

· external BG shield



OFC (5cm), Lead (15cm)	,
Boron (5cm)	neutron
Polyethylene(15cm)	neutron
EVOH sheet	radon

4 shield with door  
Easy access to chamber





### 3. Vertex/Energy reconstruction

**GEANT simulation**      **PMT hit-map**

$F(x,y,z,i)$  : acceptance of  $i^{\text{th}}$  PMT view from position  $(x,y,z)$

Interpolation with Event simulation @ 2.5[cm] lattice points of 100kg chamber



Find vertex and Energy which gives MAX Likelihood

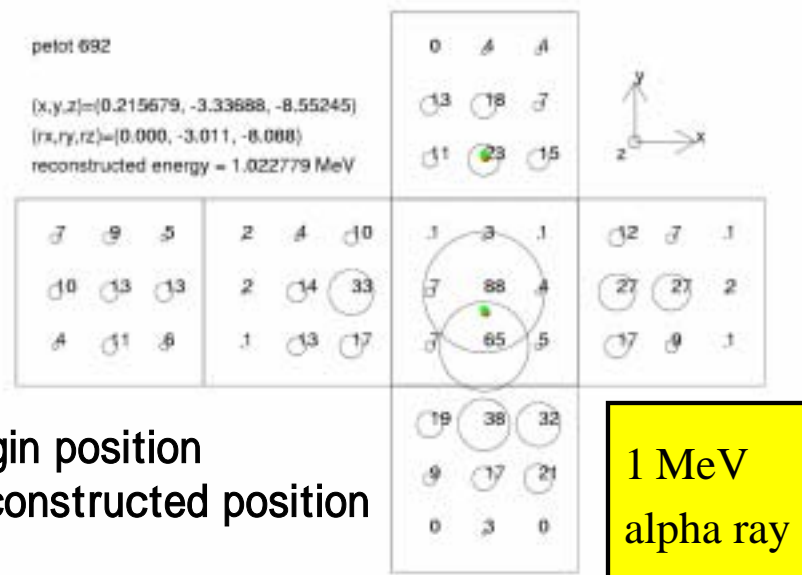
$$\text{Log}(L) = \sum_{PMT} \text{Log}\left(\frac{\exp(-\mu)\mu^n}{n!}\right)$$

L: likelihood

$\mu$ :  $F(x, y, z) \times$   
(total p.e./total acceptance)

n: observed number of p.e.

100kg XMASS detector



## 4. PMT gain calibration

- 175[nm] wave length
- compact

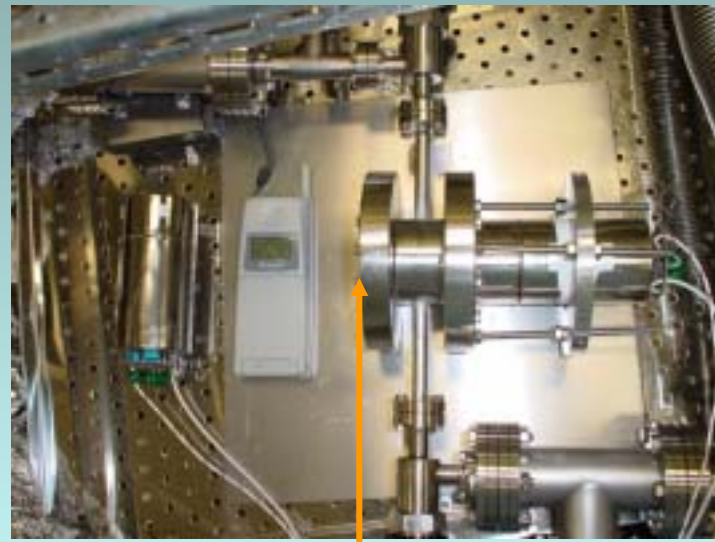


VUV standard light source

GAS-Xe chamber

+

$^{241}\text{Am}$  -source



·  $^{241}\text{Am}$  (5.45MeV)

· GAS-Xe : 2[atm]

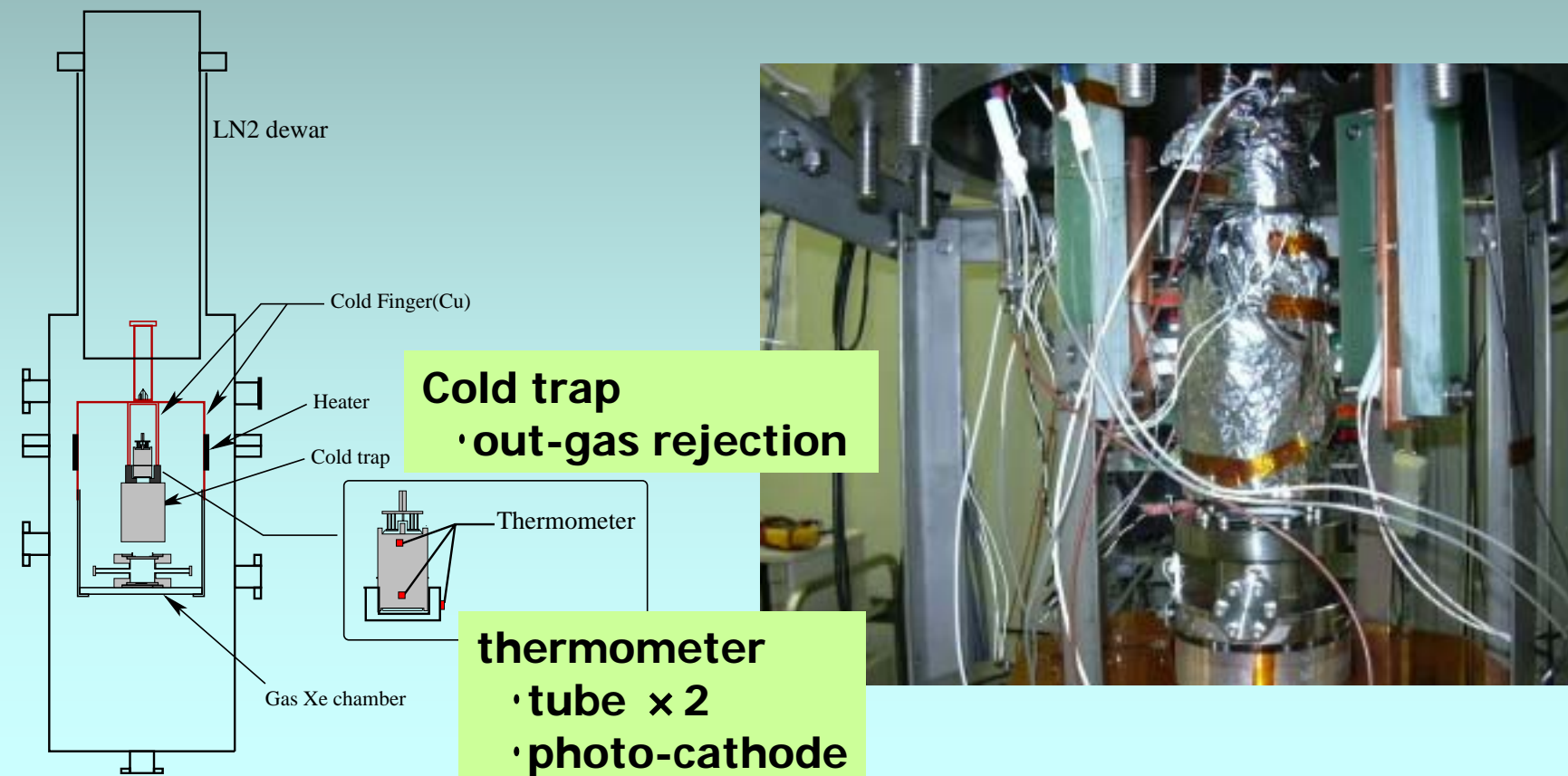
·  $\text{MgF}_2$  window : 90% transmittance@175nm

**54 PMTs were calibrated within 2%(@room temp.)**

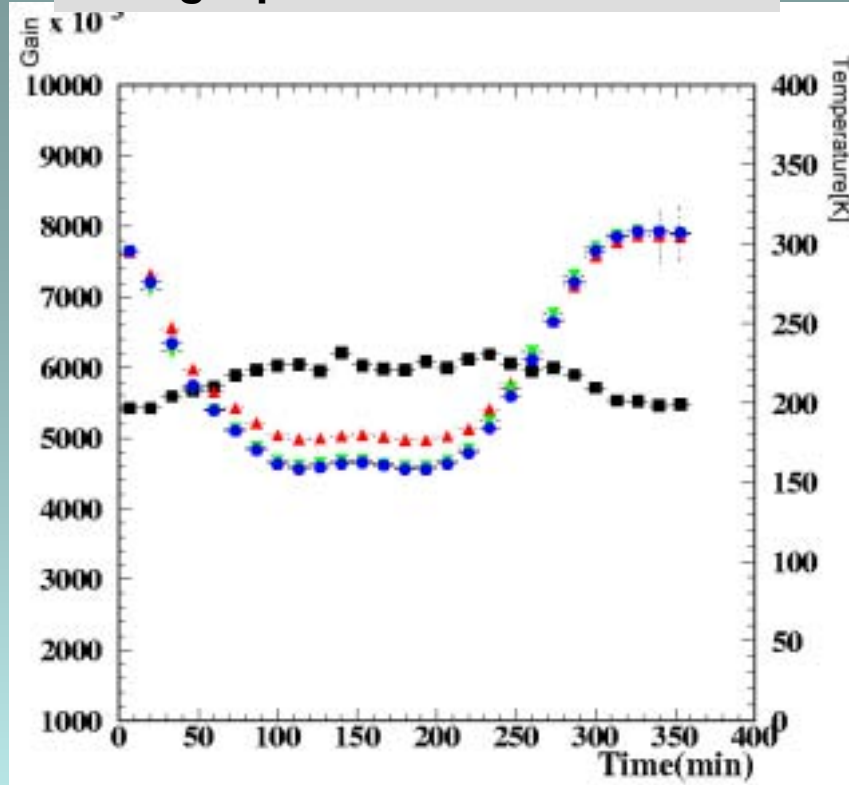
## 5. PMT cooling test

Actually, PMT temperature is about 200[K] during measurement.

multi photons measurement with gas Xe chamber : Q.E × Gain  
single photon measurement with LED : Gain

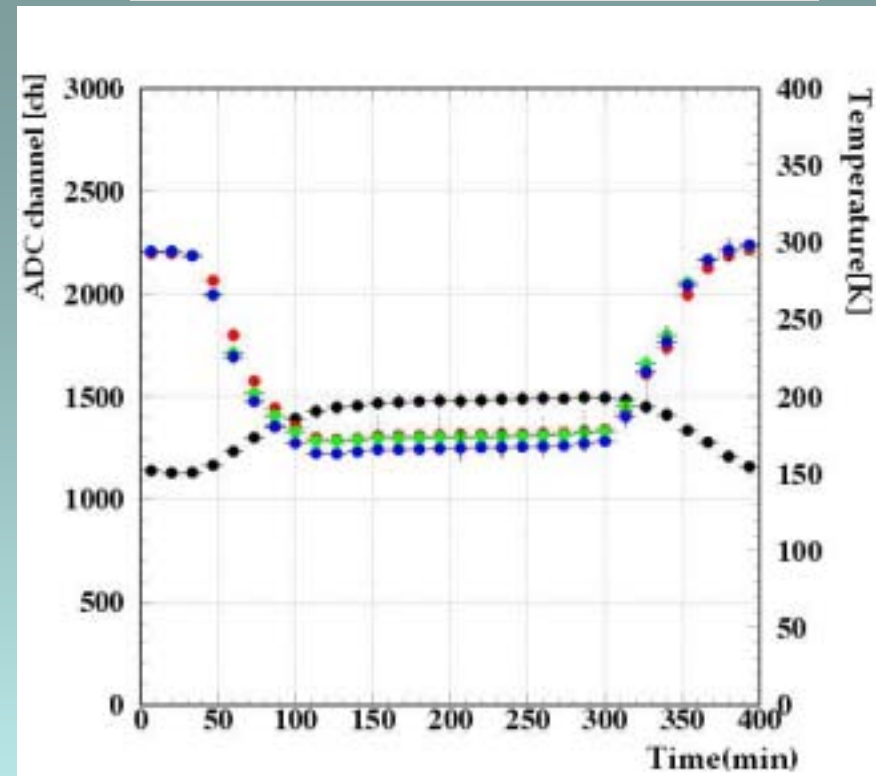


single photon measurement



Gain increase  
at low temperature

multi photons measurement



Q.E × gain also increase  
at low temperature

Measured about 4 sample PMTs

Gain increase ratio

Average : 13.9% RMS: 5.4%

Q.E increase ratio

Average : 12.2% RMS : 4.7%

There is individual difference

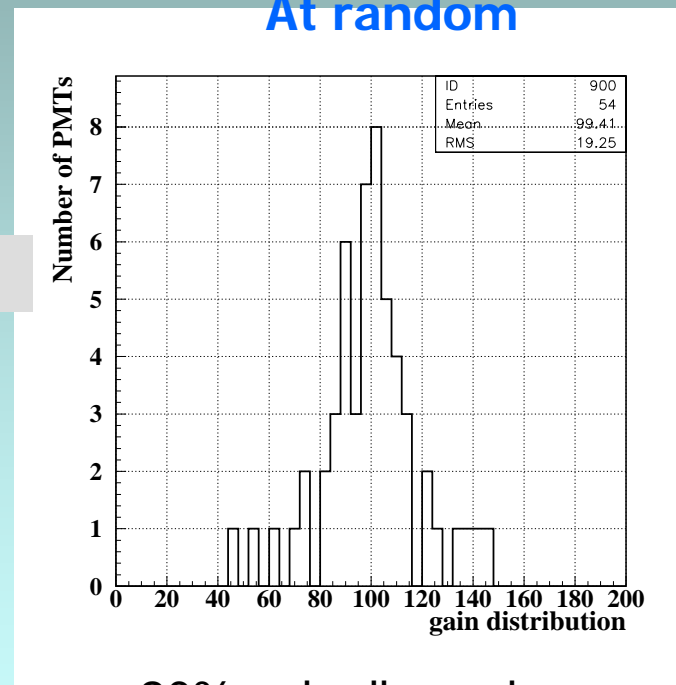
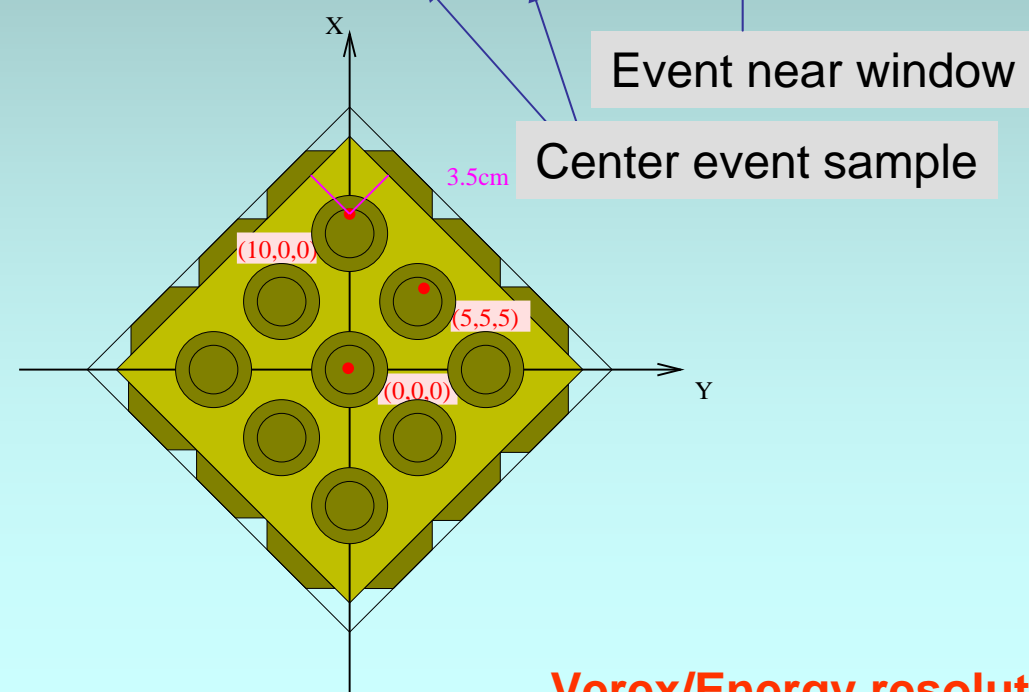
## assumption of gain calibration from vertex/Energy reconstruction simulation

EVENT simulation by GEANT  
get p.e. MAP about 54 PMTs



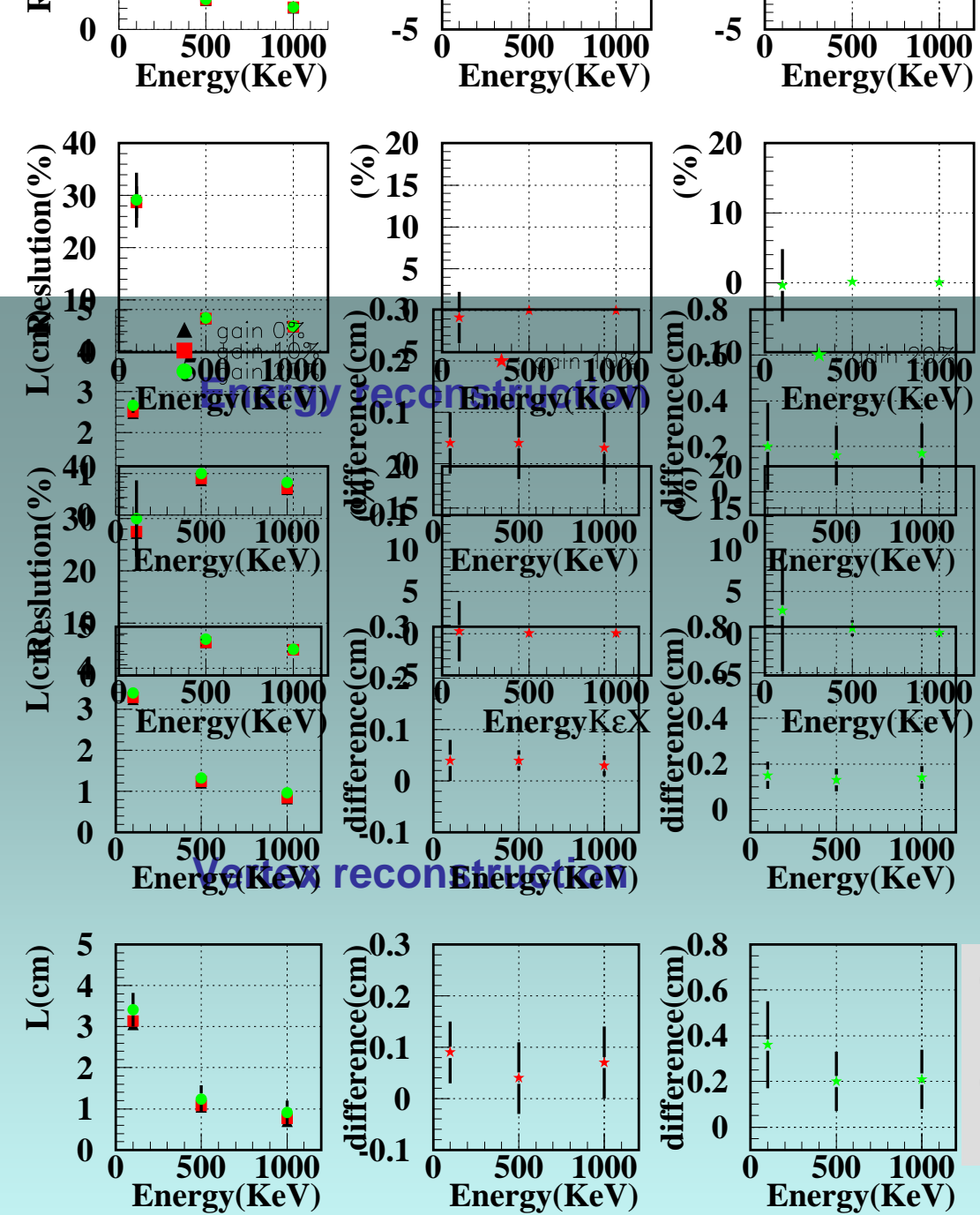
Input  
10% and 20% Gain Dispersion  
At random

Energy : 100keV ,500keV, 1MeV  
Position : (0,0,0) (5,5,5) (10,0,0)



20% gain dispersion

Vertex/Energy resolution ??



Gain : 20%  
Energy resolution < 10%

No problem

Gain : 20%  
position resolution > 0.5cm  
Gain : 10%  
position resolution < 0.15cm

Event reconstruction request for gain within less than 10% accuracy

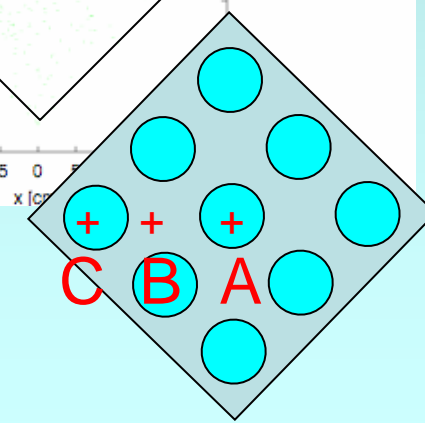
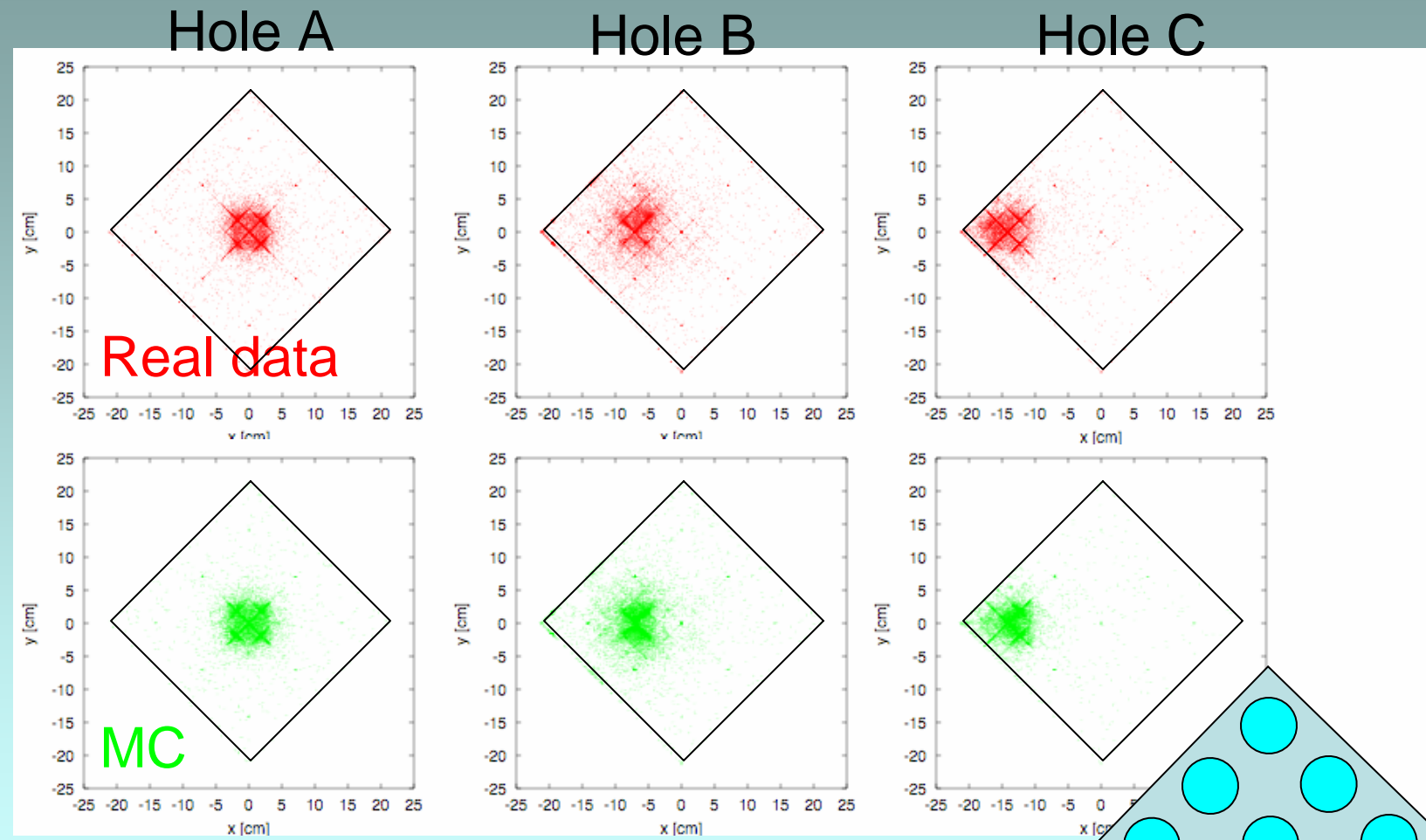
Dispersion of PMT gain at low temperature is no problem

## **6. Data analysis**

**vertex/Energy reconstruction  
Demonstration of self-shielding  
Low Background**

# Self shielding performance

Scatter plot of 3 collimators' run

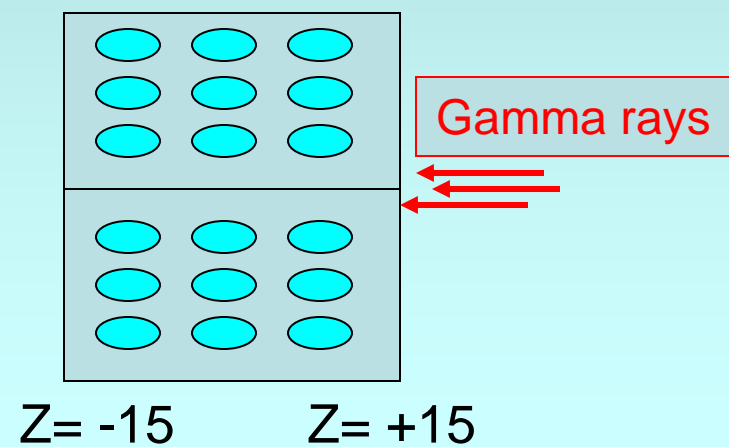
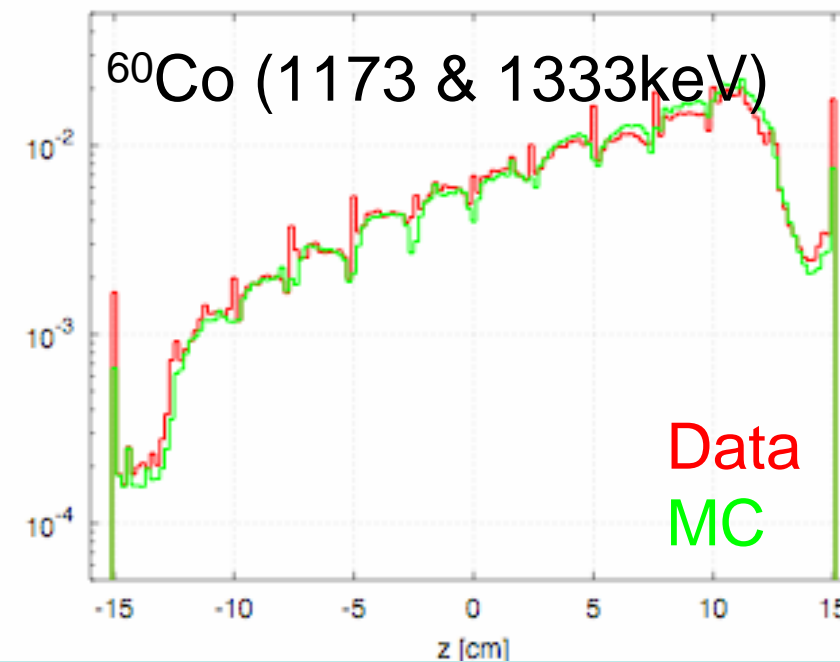
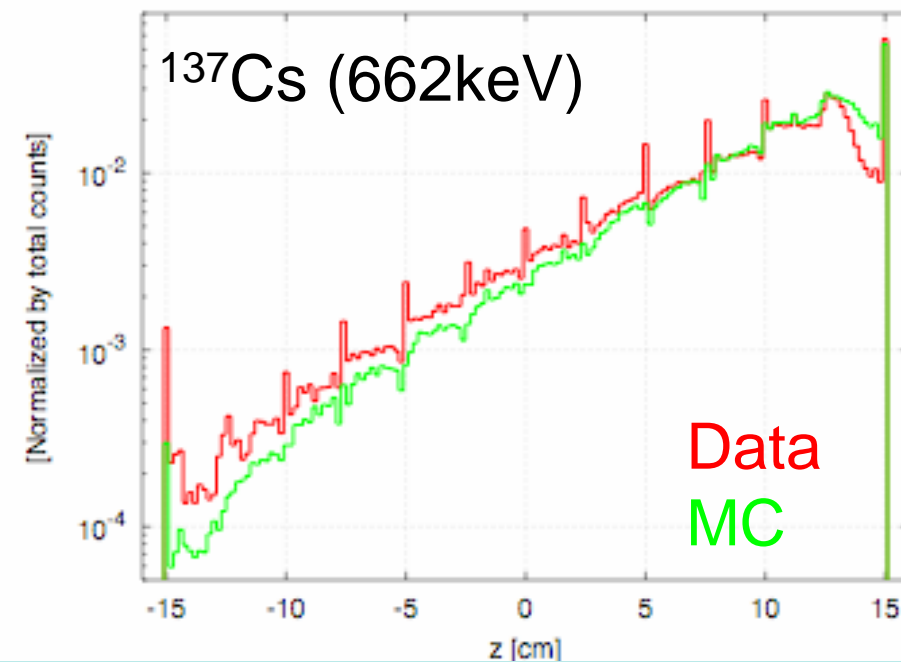


same shape!!



# Self shielding performance

Z position distribution of photoelectric peaks

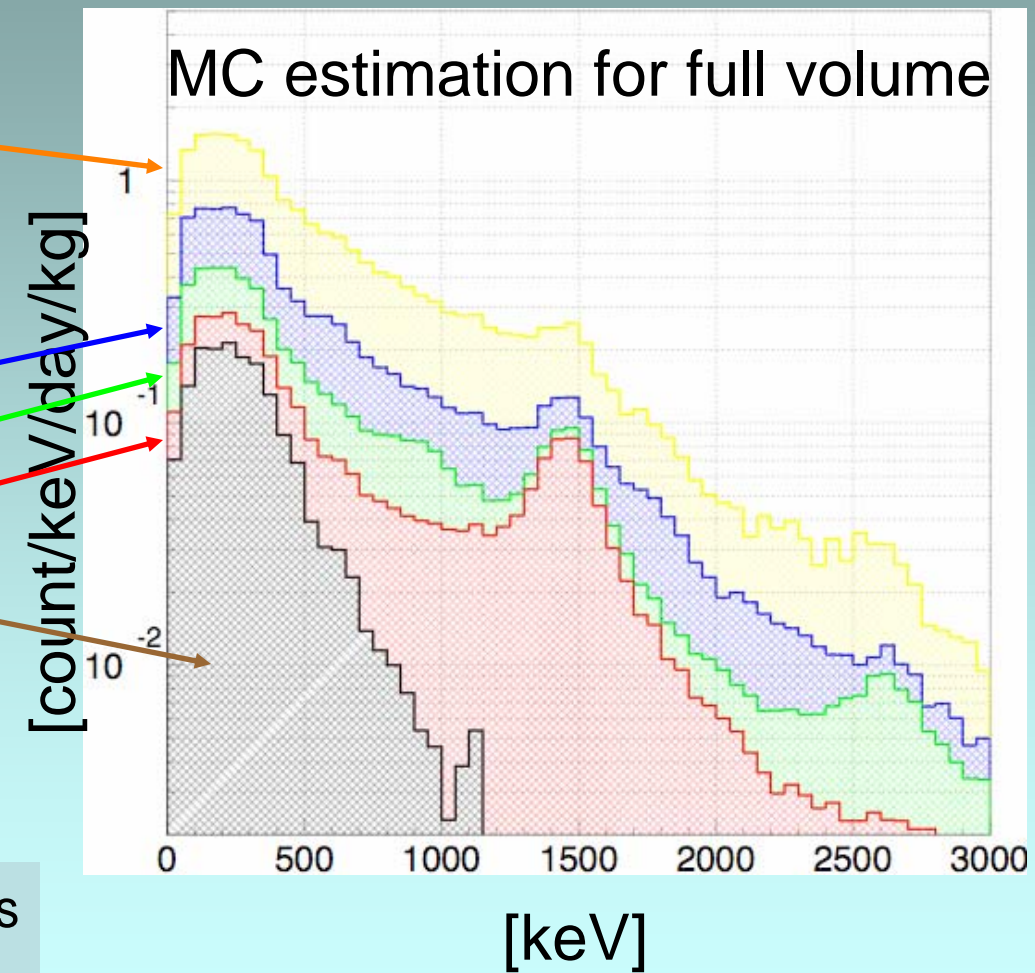


- Good agreement with MC
- Self shielding power works well as expected
- (Need to improve fitter)

# Background estimated from known sources

- From outside of the shield:  
0.71 g/cm<sup>2</sup> (>500keV)
- RI sources inside of the shield
  - PMTs (Bq/PMT)
    - <sup>238</sup>U :  $1.8 \times 10^{-2}$
    - <sup>232</sup>Th:  $6.9 \times 10^{-3}$
    - <sup>40</sup>K :  $1.4 \times 10^{-1}$
- Pb-214 in the lead shield:  
250 Bq/kg

- ◆ Xenon internal radioactivities
  - <sup>238</sup>U, <sup>232</sup>Th contamination
  - <sup>85</sup>Kr contamination

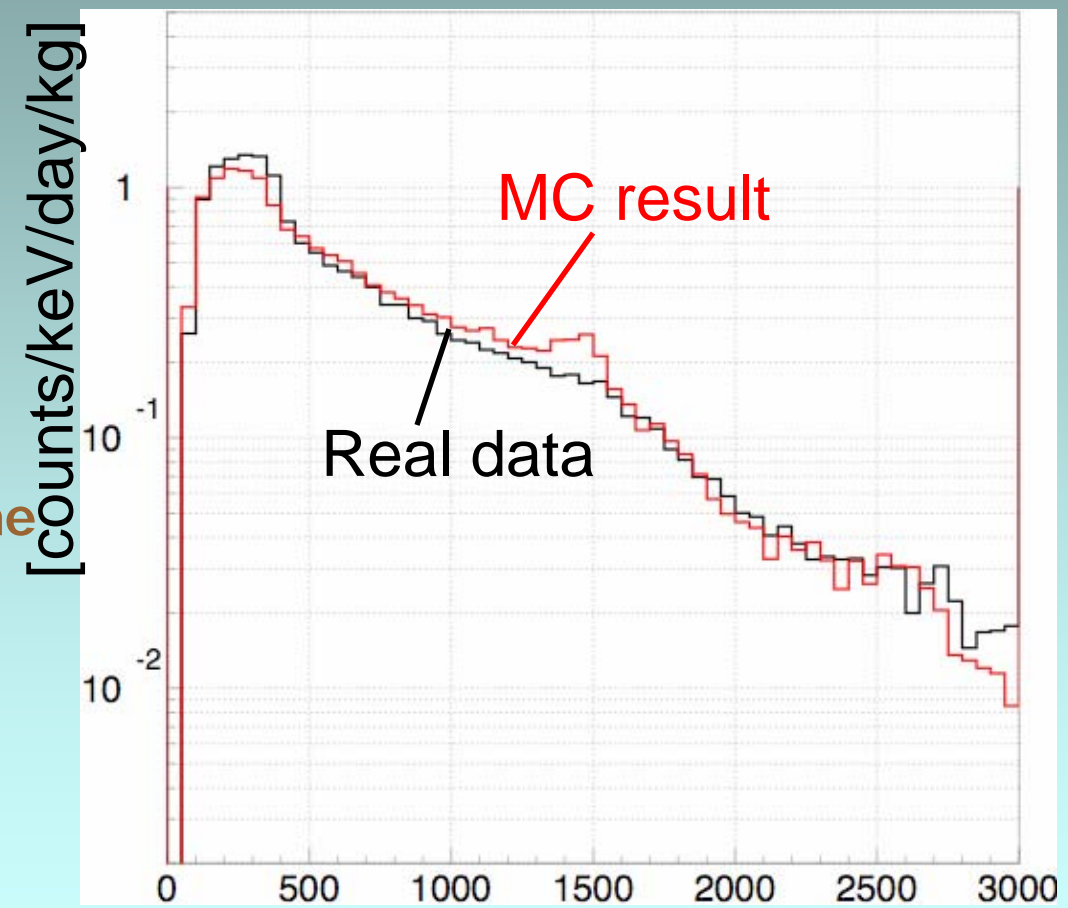


# BG comparison with MC estimation

- **MC**  
Gamma from outside  
RI in PMTs(U,Th,K)  
 $^{210}\text{Pb}$  in the lead shield

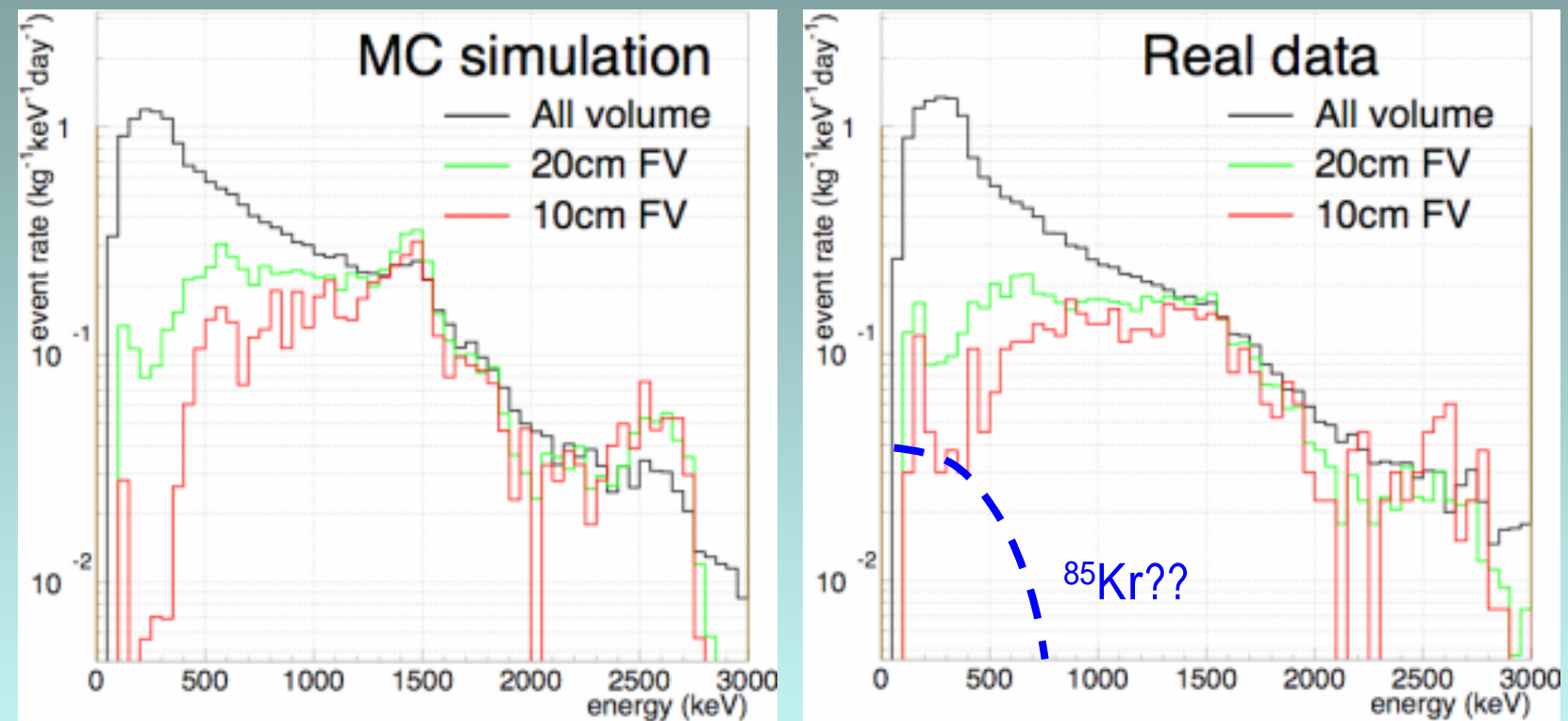


- **Real data for full volume**  
(lifetime: 0.9days)



**Good agreement!** [keV]

## Self shielding power

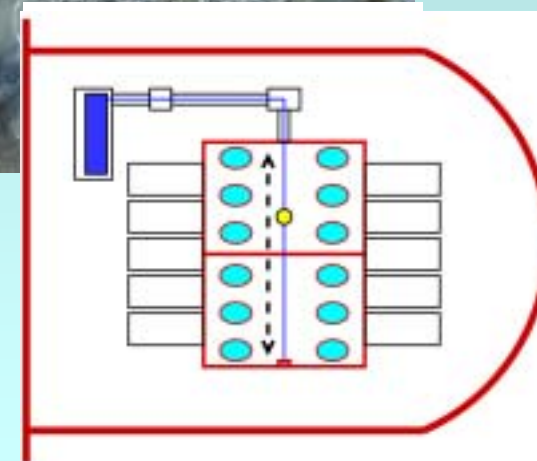


- Background decreases with the fiducial volume cut  
     $\longrightarrow$  **Ultra-low BG** at the inner volume
- Something exists at low energy region  
     $\longrightarrow$   $^{85}\text{Kr}??$

Will be compared after processed by the distiller

# Future plan with 100kg detector

- 1st run: **DONE!**
  - Confirmed the basic properties
  - Evaluated the event reconstruction performance
  - Background measurements and breakdown of its origin
- **Further study:**
  - Detailed study of event reconstruction  
Source run with inner sources
  - Detailed study of the background  
With a distiller and various purification systems  
With a neutron source



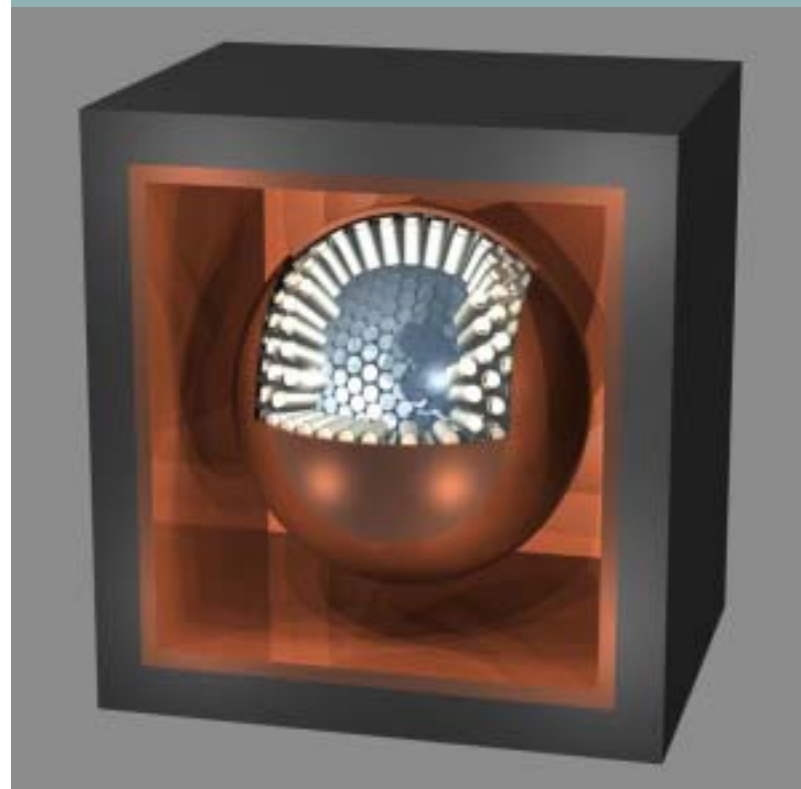
➔ Development of the large size detector

# **800kg Detector simulation/future plan**

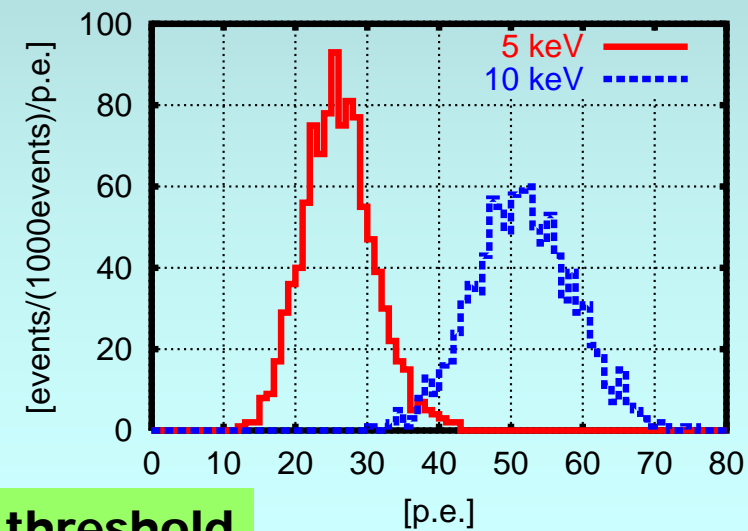
- 1. Introduction**
- 2. BG simulation**
- 3. Expected Sensitivity for DM**

# 1. Introduction of 800kg detector

## DARK MATTER search



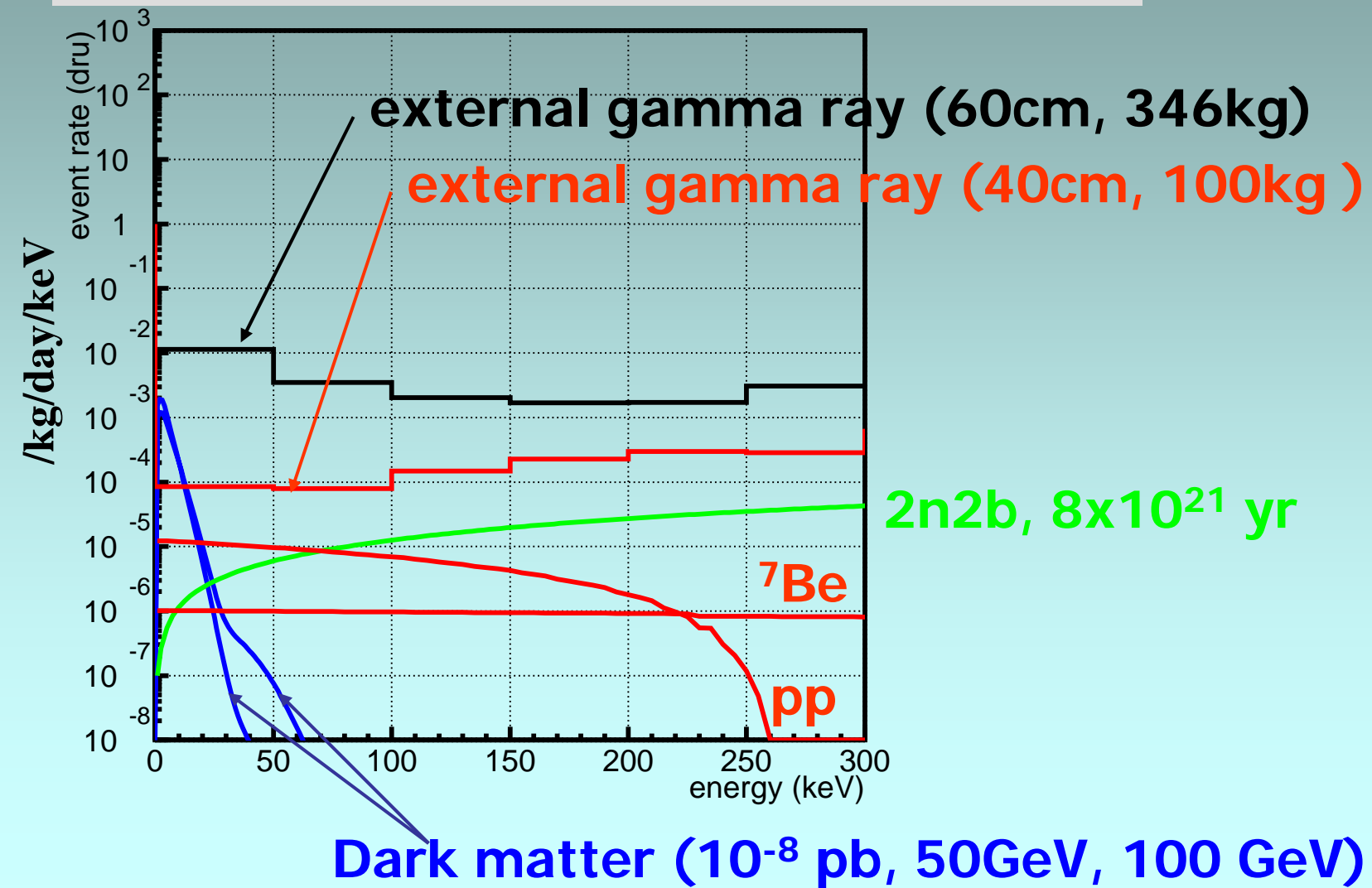
100kg liquid Xe  
~80cm diameter sphere  
About 640 2-in PMTs  
75% photo-coverage  
5 p.e./keV



Very low Energy threshold

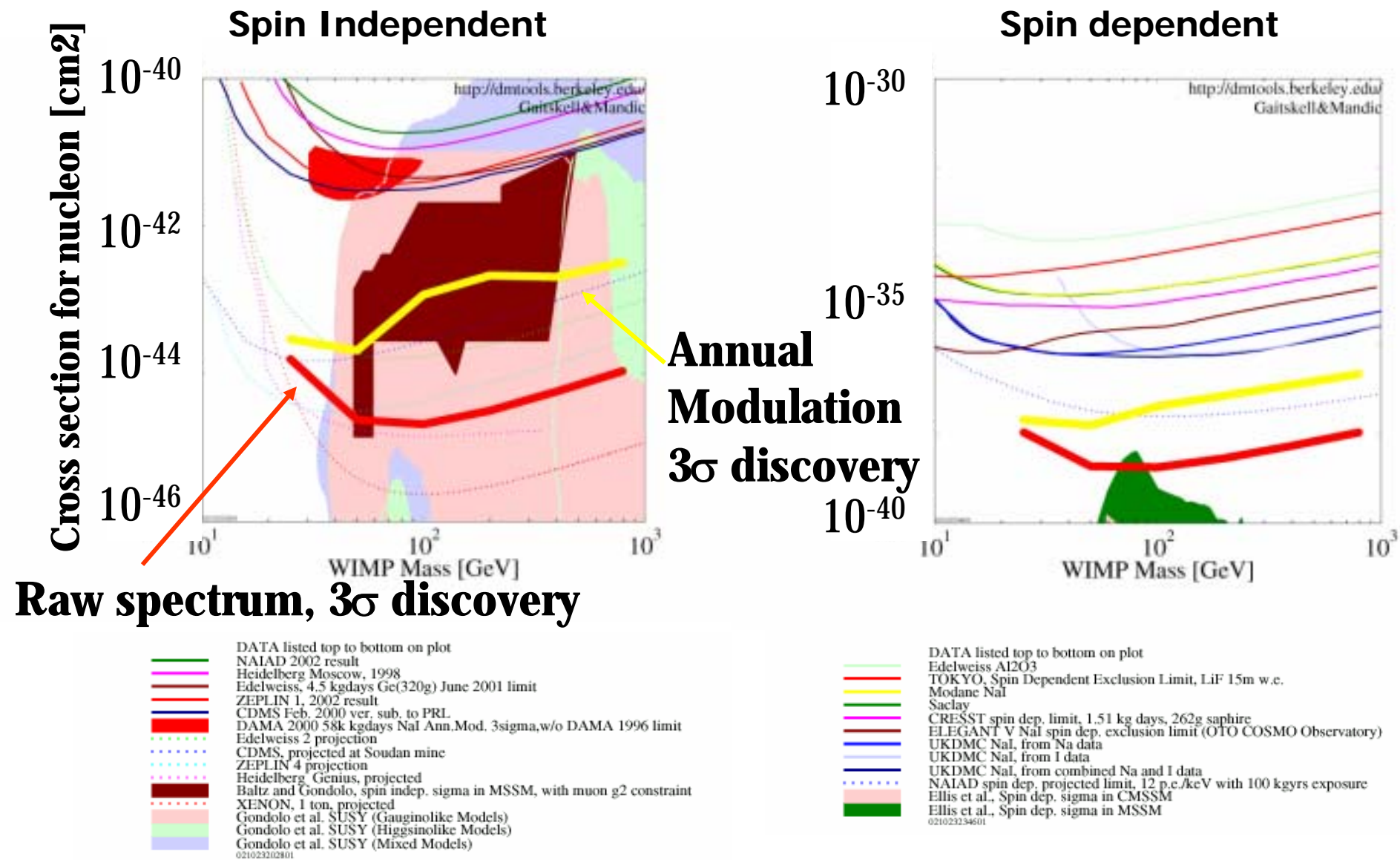
## 2. External background in 800kg detector

- Dominant contribution is from PMT
- Assuming further 1/10 reduction of PMTs BG





### 3. Expected sensitivity for DM



# Summary

## **XMASS experiment:**

Ultra low background experiment with liquid Xenon  
And there are some physical purposes.

### **1st run of 100kg detector was done:**

**Event reconstruction worked well**  
**Background level was low as expected**  
**Self-shielding power was confirmed**

### **Next 800kg detector:**

**Designed for dark matter search**  
**Will has a extremely high sensitivity for DM detection**